

GRONINGEN, THE NETHERLANDS
APRIL 17-18, 2019



university of
 groningen

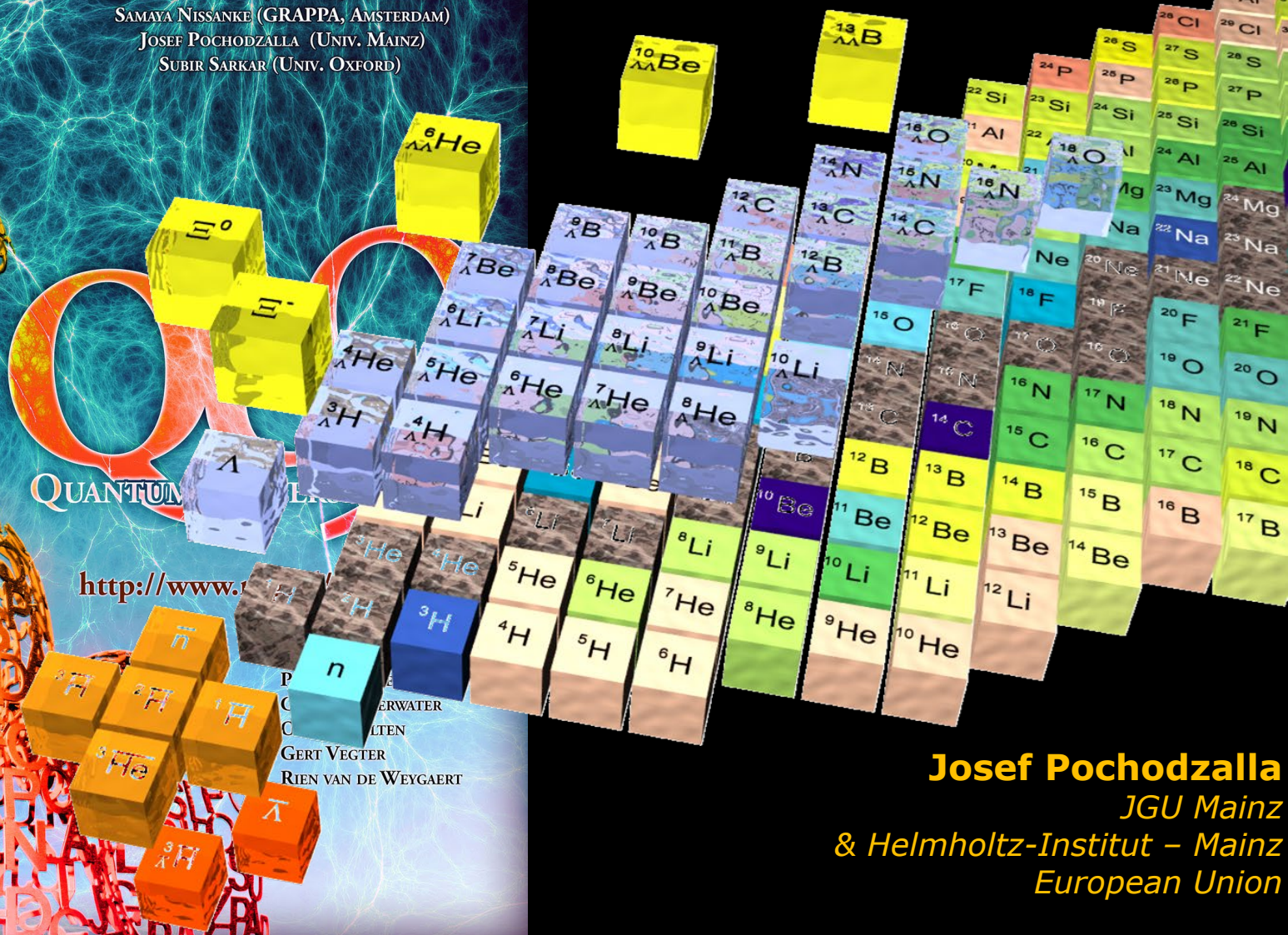
INVITED SPEAKERS:

- SYTZE BRANDENBURG (KVI-CART, GRONINGEN)
- JULIA EVEN (KVI-CART, GRONINGEN)
- JACQUES LASKAR (CNRS, OBSERV. PARIS)
- THOMAS MORGAN (DIFFER, EINDHOVEN)
- SAMAYA NISSANKE (GRAPPA, AMSTERDAM)
- JOSEF POCHODZALLA (UNIV. MAINZ)
- SUBIR SARKAR (UNIV. OXFORD)

QU9 – Quantum Universe Groningen - 18 April 2019

Hypernuclei & Hyperatoms

Bringing heaven to earth



QUANTUM

<http://www.>

PROF. DR. GERT VEGTER
RIEN VAN DE WEYGAERT

Josef Pochodzalla
JGU Mainz
& Helmholtz-Institut – Mainz
European Union

First multi-messenger observations of a binary neutron star merger



Neutron stars are Superstars

super high density
super strong magnetic fields
super fast rotation
super strong gravity *in Matter*

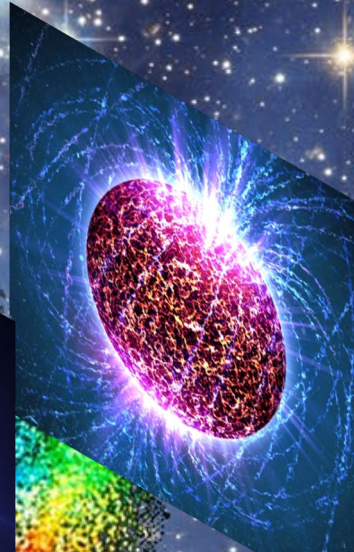
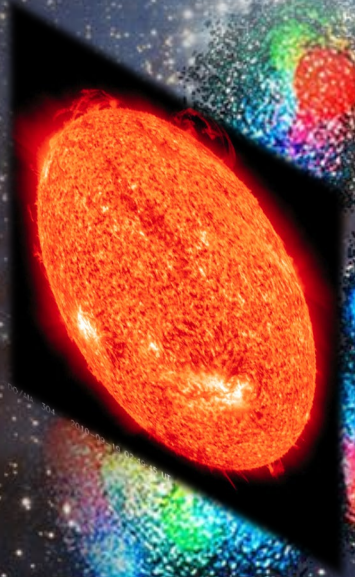
$$1 - \frac{2GM}{c^2 R}$$

~ 0.3

$\sim 10^{-4}$

$\sim 10^{-7}$

$\sim 10^{-10}$

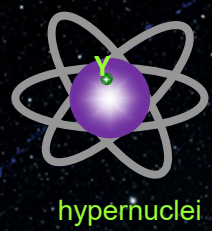
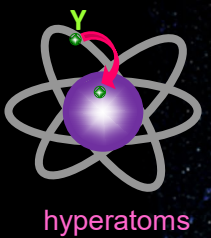


~ 1 million
black holes

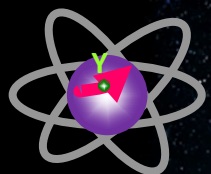
~ 100 million
neutron stars

~ 10 billion
white dwarfs

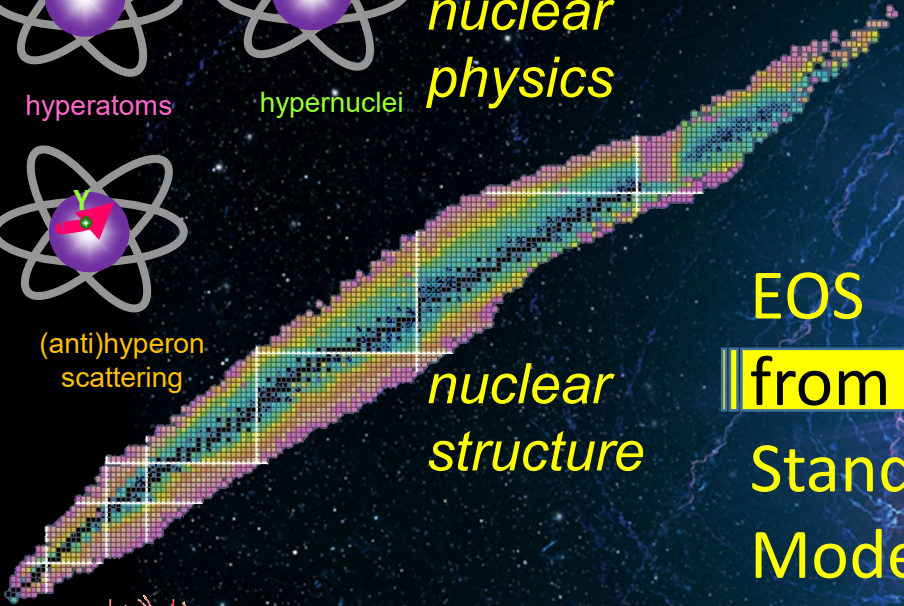
in our galaxy ~ 300 billion stars



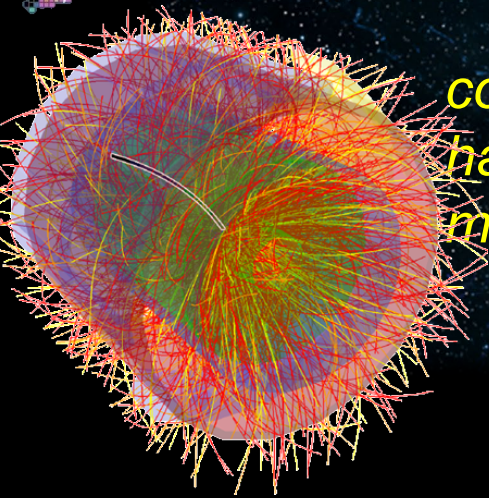
strangeness
nuclear
physics



(anti)hyperon
scattering

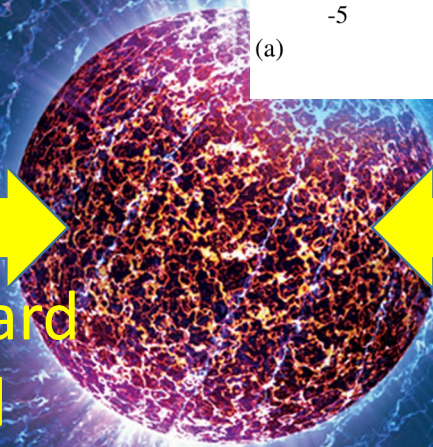


nuclear
structure

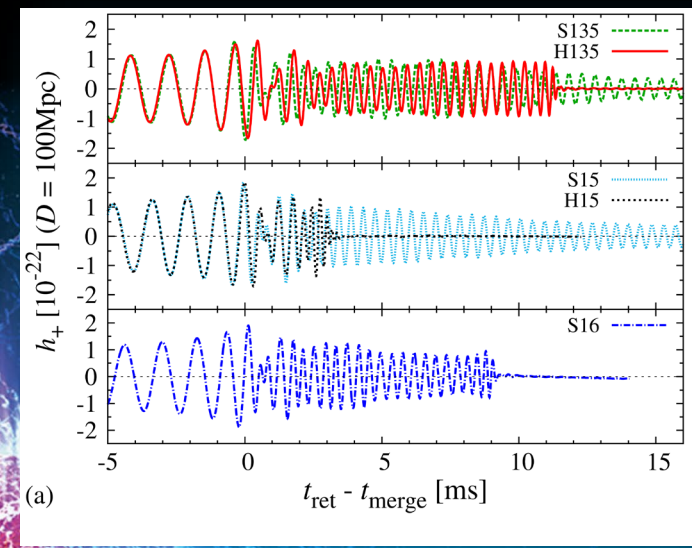


compressed
hadronic
matter

EOS
from
Standard
Model



EOS
from
Standard Model
+strong field
GRAVITY



NEUTRON STAR MODELS

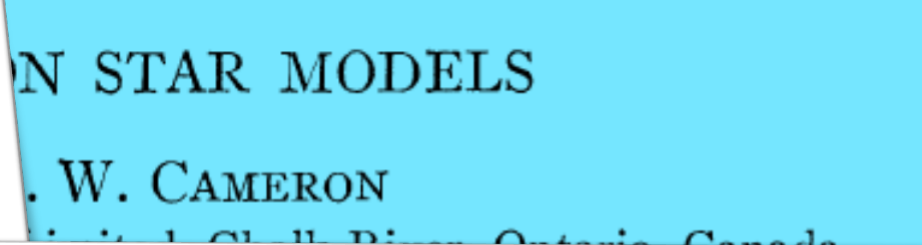
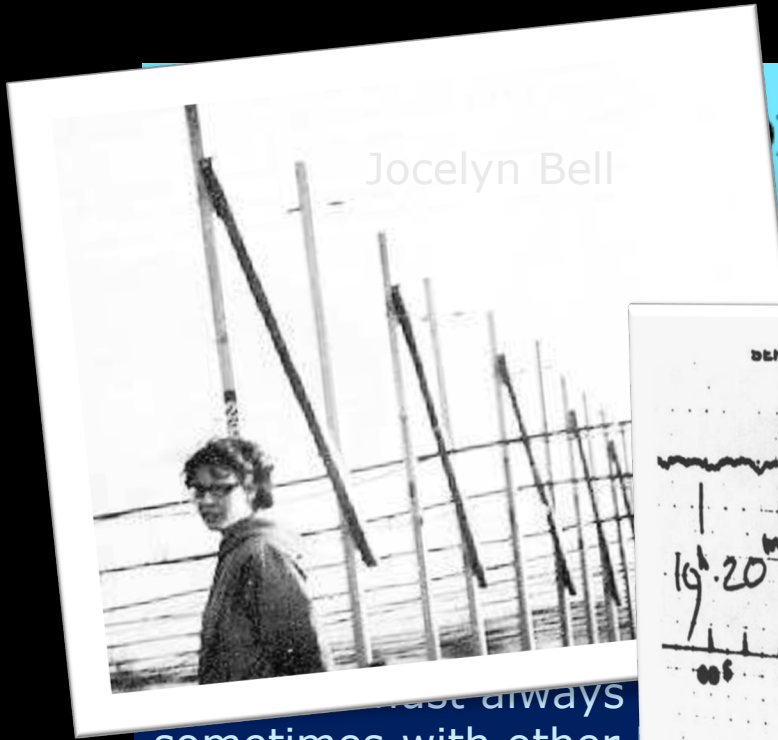
A. G. W. CAMERON

Atomic Energy of Canada Limited, Chalk River, Ontario, Canada

Received June 17, 1959

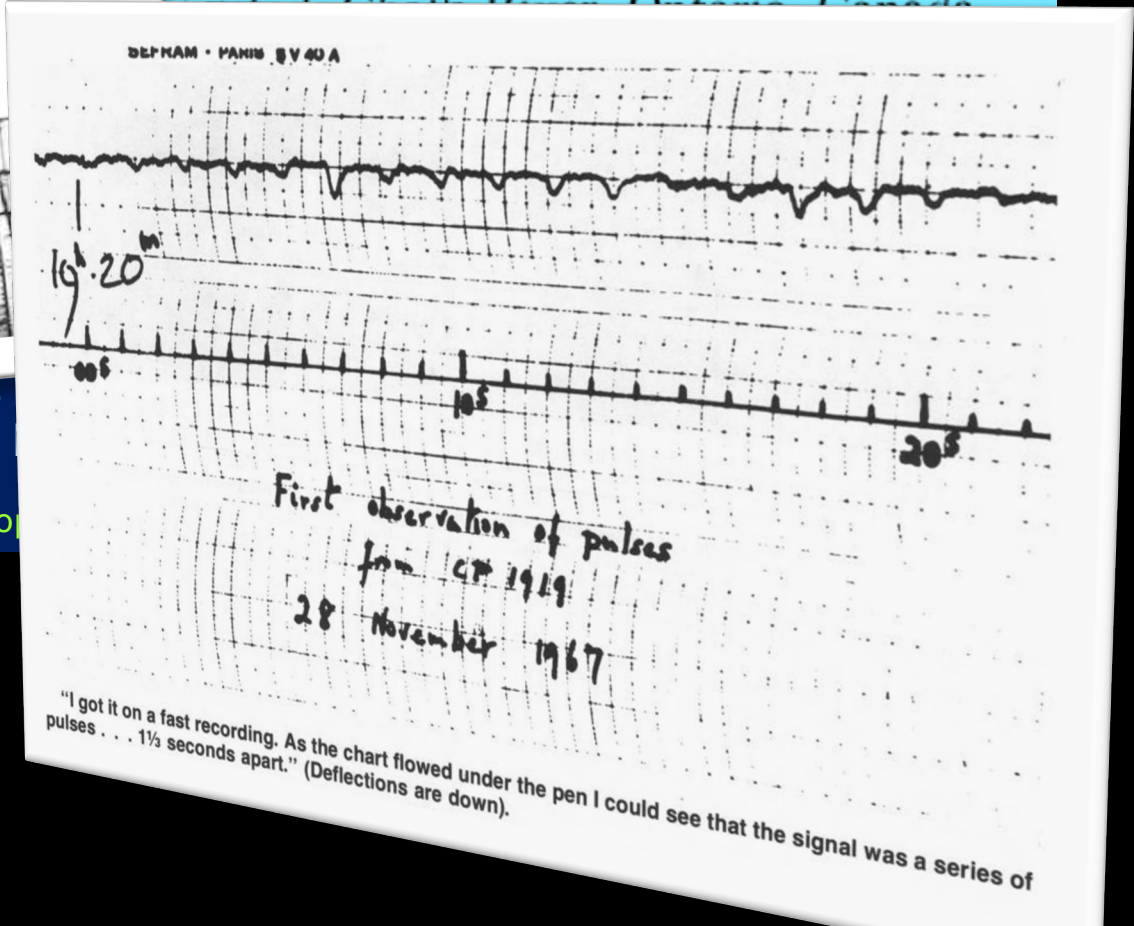
„Another reason why the writer has not taken into account complications inherent in using a realistic equation of state is that no such things such pure neutron stars can be expected to exist. The neutrons must always be contaminated with some protons and sometimes with other kinds of nucleons (hyperons or heavy mesons).„

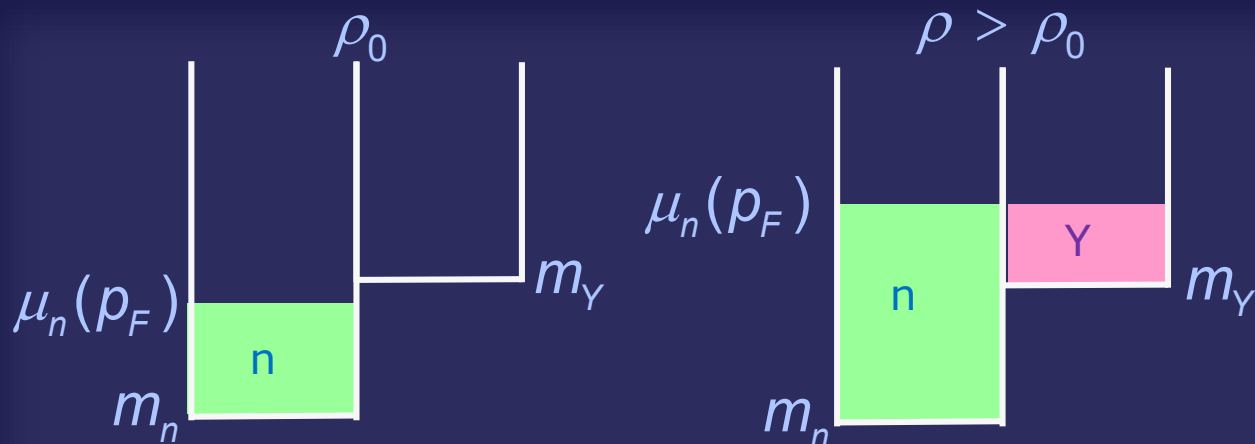
Alastair G.W. Cameron, *Astrophysical Journal*, vol. 130, p.884 (1959)



... always
sometimes with other

Alastair G.W. Cameron, *Astro*





$$p_{F,n}^2 + m_n^2 \geq m_\Lambda^2$$

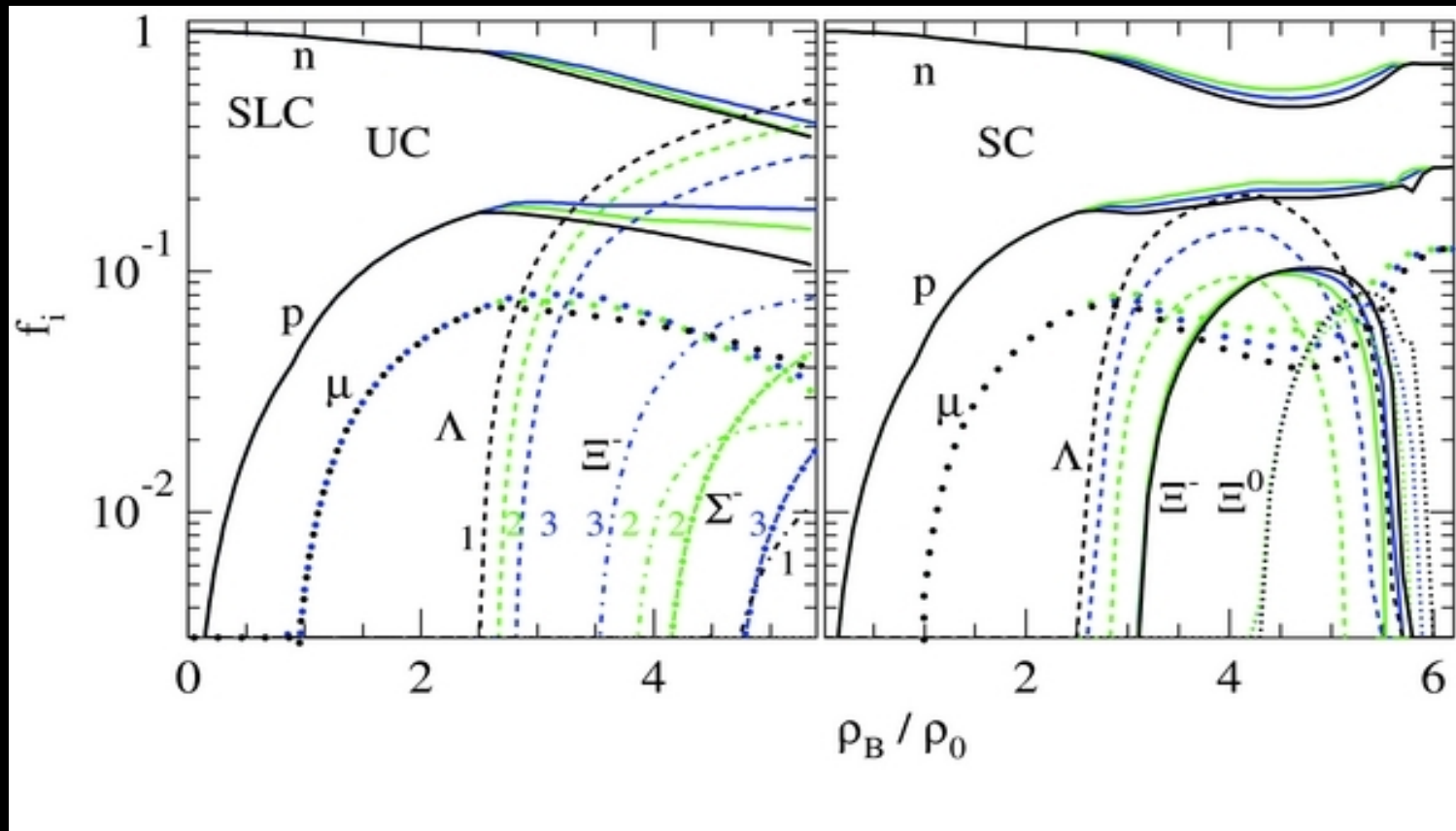
$$m_\Lambda = 1116 \text{ MeV}, m_n = 939 \text{ MeV} \Rightarrow p_{F,n} \approx 600 \text{ MeV} \approx 3 \text{ fm}^{-1}$$

$$\text{non-interacting Fermi-gas: } \rho = \frac{p_{F,n}^3}{3\pi^2} \Rightarrow p_{F,n}(\rho_0) = 1.7 \text{ fm}^{-1}$$

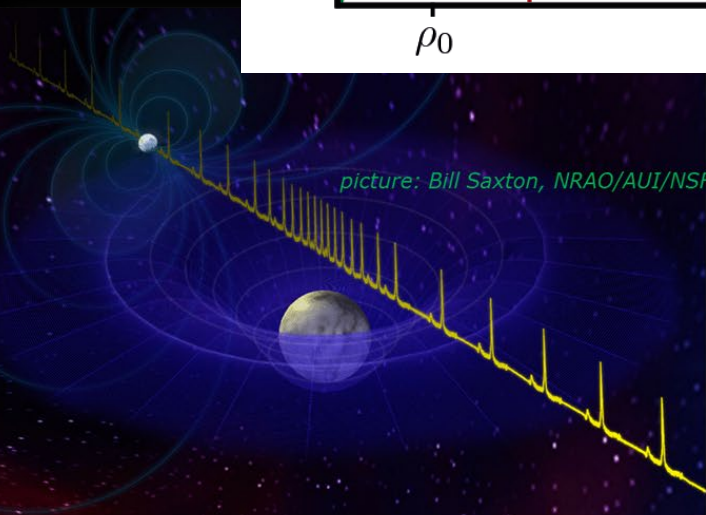
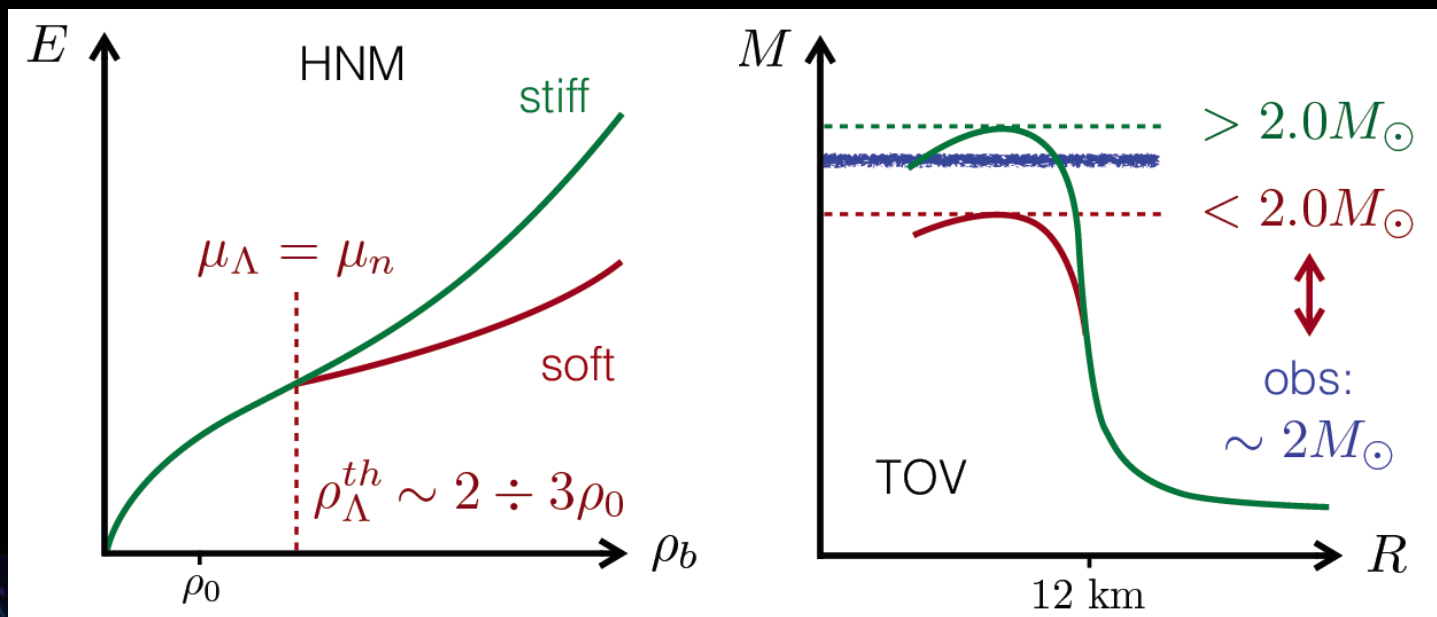
\Rightarrow appearance of hyperons at $\rho_\Lambda \approx 5.5\rho_0$

with interactions $\rho_\Lambda \approx 2 - 3\rho_0$

- Sequence of hyperon appearance depends on B-B interaction
- Σ -N interaction repulsive \Rightarrow Σ will probably appear latest



appearance of hyperons \Rightarrow relieve of Fermi pressure \Rightarrow softer EOS
 \Rightarrow reduction of maximal mass



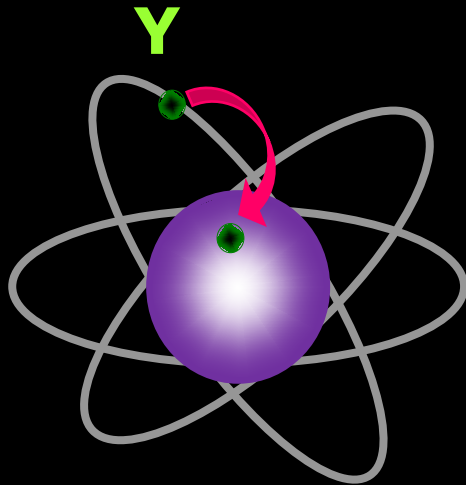
$$M(\text{PSR J1614-2230}) = 1.928 \pm 0.017 M_\odot$$

$$M(\text{PSR J0348+0432}) = 2.01 \pm 0.04 M_\odot$$

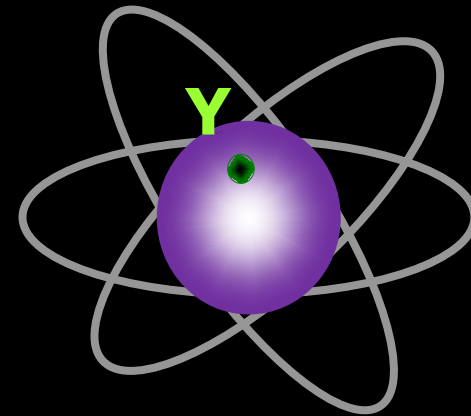
$$M(\text{PSR J1946+3417}) = 1.828 \pm 0.022 M_\odot$$

P. B. Demorest *et al.*, Nature 467 (2010)
 update: E. Fonseca *et al.*, ApJ 832, 167 (2016)
 J. Antoniadis *et al.*, Science 340 (2013)
 E.D. Barr *et al.*, MNRAS 465, 1711–1719 (2017)

hyperatoms



hypernuclei

**Objects**

Hyperons in atomic levels
within the nuclear periphery

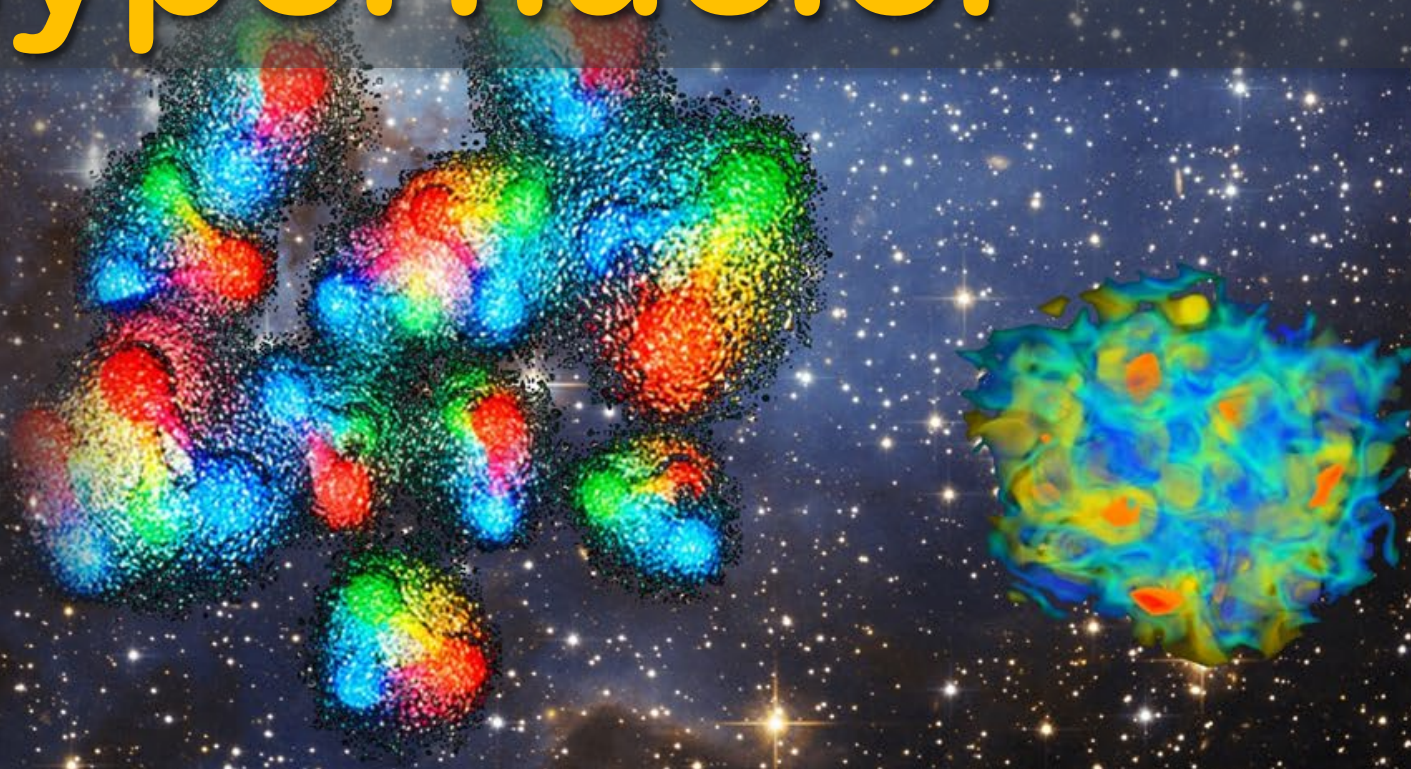
Hyperons bound by strong
interaction within nuclei

Methodology

Width and shift of Σ^- , Ξ^- and Ω^-
atomic levels

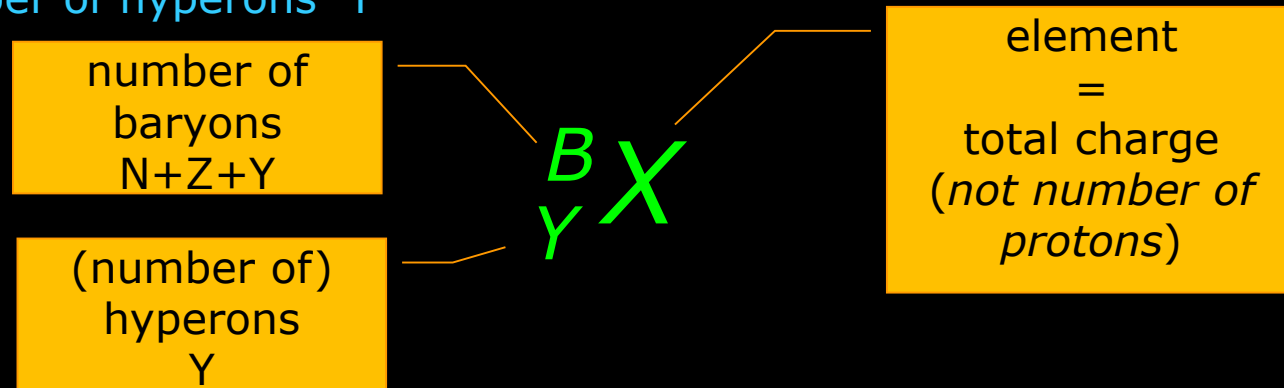
Masses, excited state
spectrum of Λ and $\Lambda\Lambda$
hypernuclei,

Hypernuclei

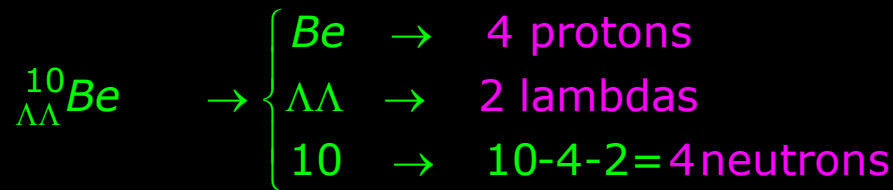
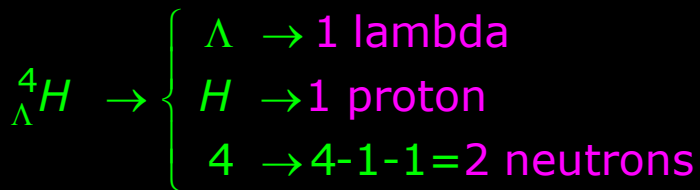


What is a hypernucleus?

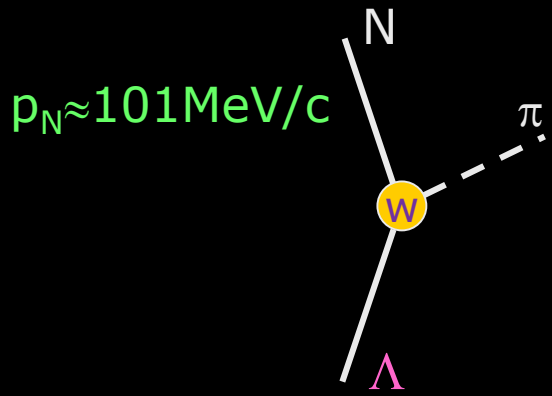
- a hypernucleus is specified by
 - ▶ the number of neutrons N
 - ▶ the number of protons Z
 - ▶ the number of hyperons Y



- ▶ since we have more than one hyperon (Λ , Ξ^- , Σ^{+0}) one usually writes explicitly the symbols of one (or more) hyperon
- ▶ examples:



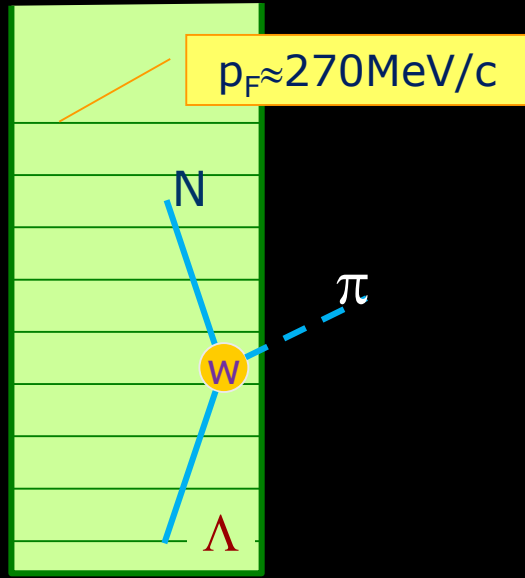
free Λ decay



$\Lambda \rightarrow p\pi^- + 38\text{MeV}$ (64%)
 $\Lambda \rightarrow n\pi^0 + 41\text{MeV}$ (36%)
 $\tau_\Lambda = 263\text{ps}$

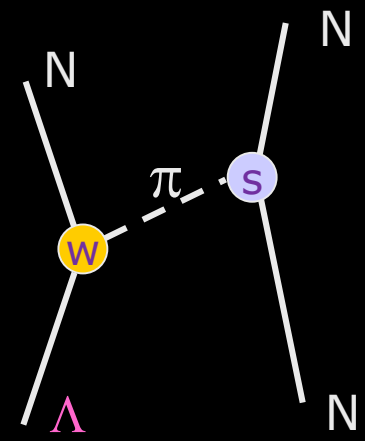
$\Delta I = 1/2$ rule

mesonic decay of hypernuclei



suppressed by Pauli blocking

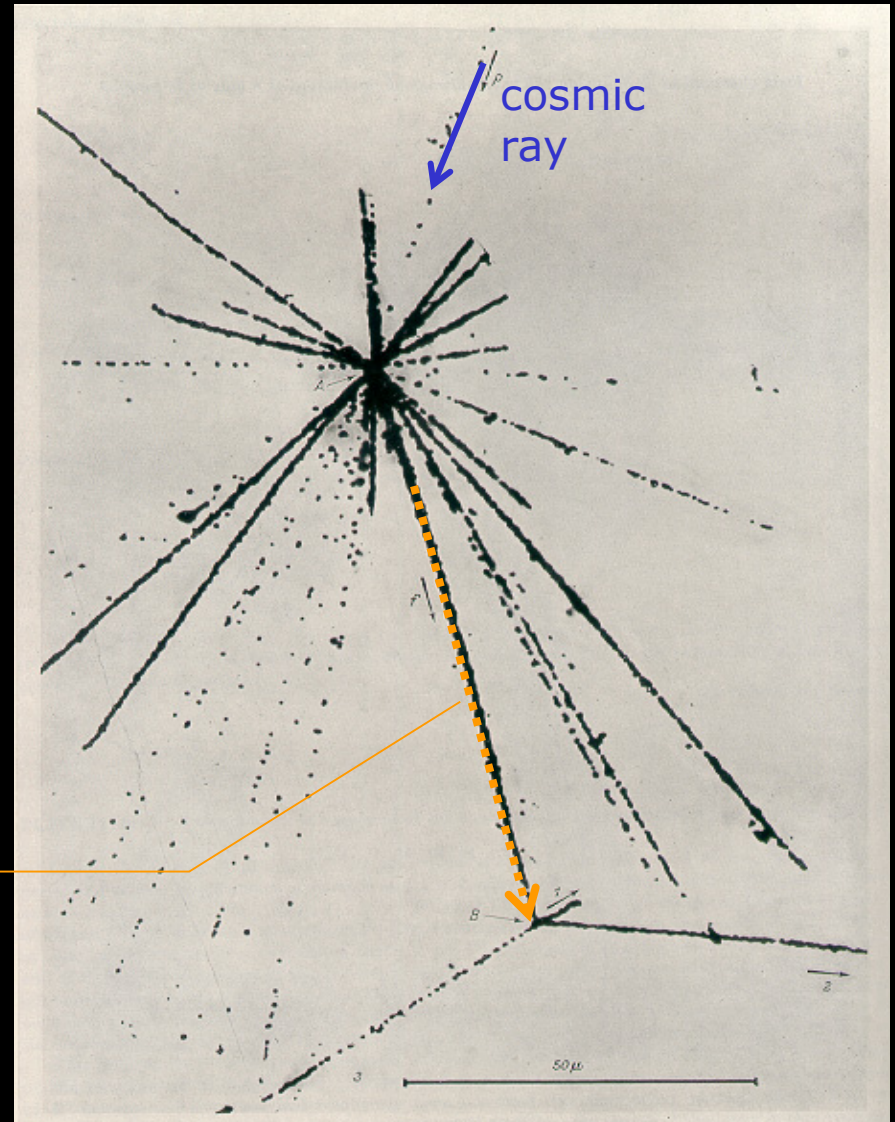
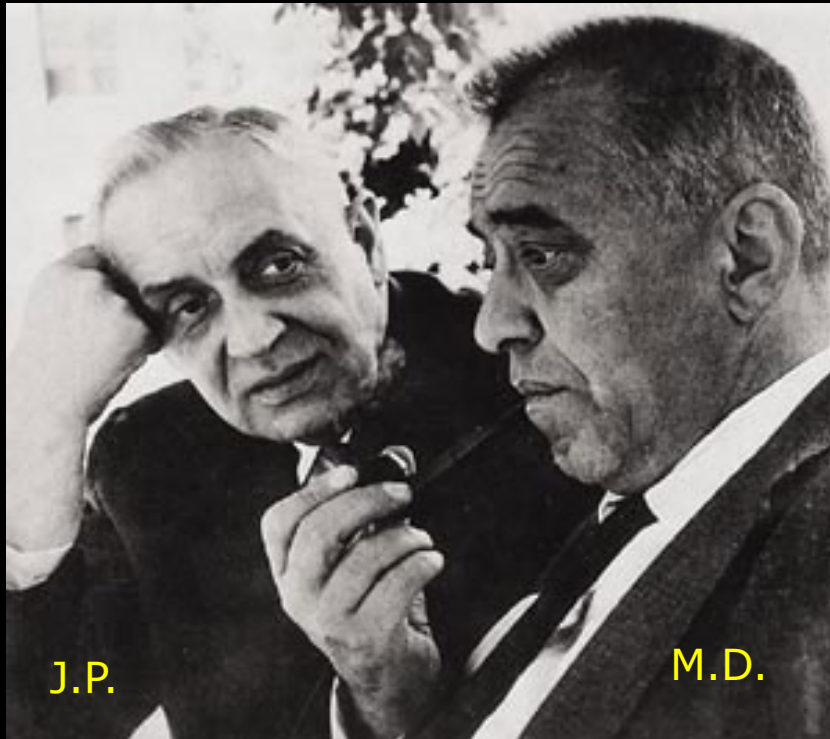
non-mesonic decay of hypernuclei



$\Lambda p \rightarrow np + 176\text{MeV}$
 $\Lambda n \rightarrow nn + 176\text{MeV}$

dominant in all but the lightest hypernuclei

- Marian Danysz, Jerzy Pniewski, et al. Bull. Acad. Pol. Sci. III **1**, 42 (1953)
- Marian Danysz, Jerzy Pniewski, Phil. Mag. **44**, 348 (1953)

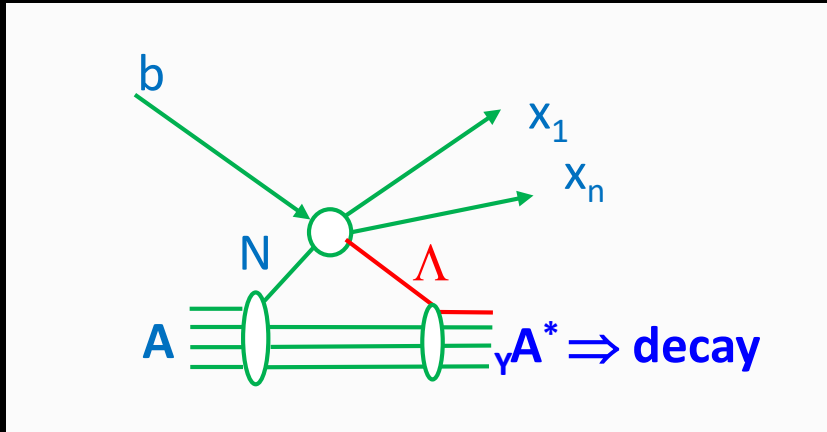
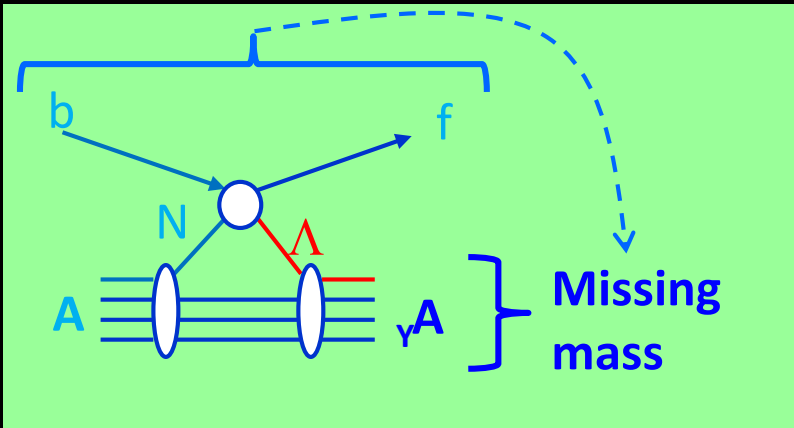


$$t > \frac{s}{c} \sim \frac{80 \mu\text{m}}{300000 \text{km/s}} \approx 2.6 \cdot 10^{-13} \text{s}$$

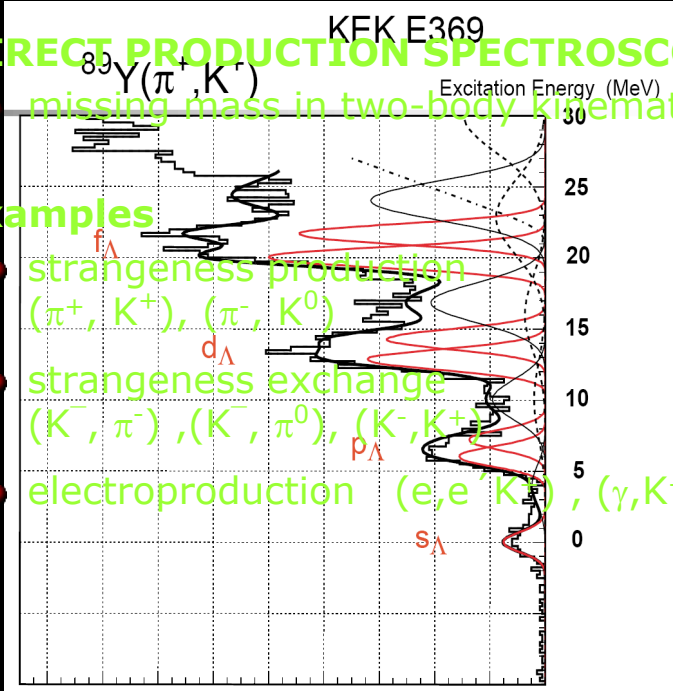
$$\tau(\Lambda) = 2.6 \cdot 10^{-10} \text{s}$$

⇒ typical for weak decay

The twofold way to hypernuclei



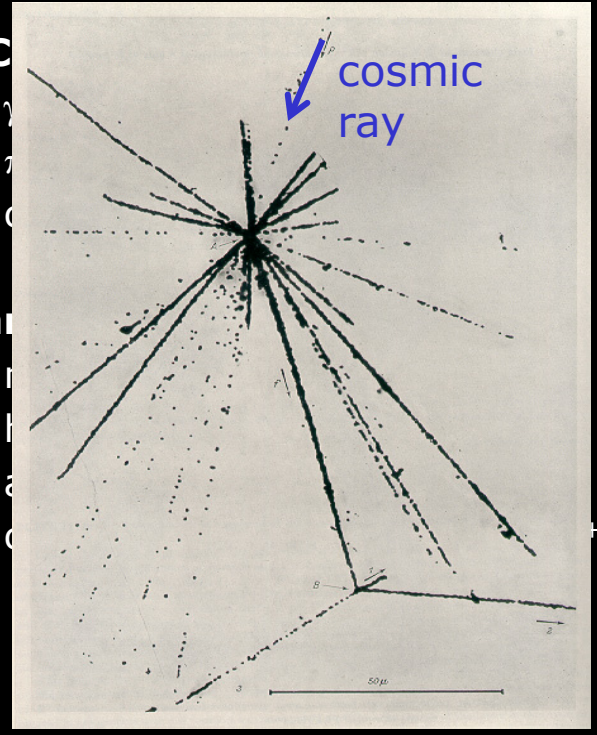
• **DIRECT PRODUCTION SPECTROSCOPY**
 • missing mass in two-body kinematics



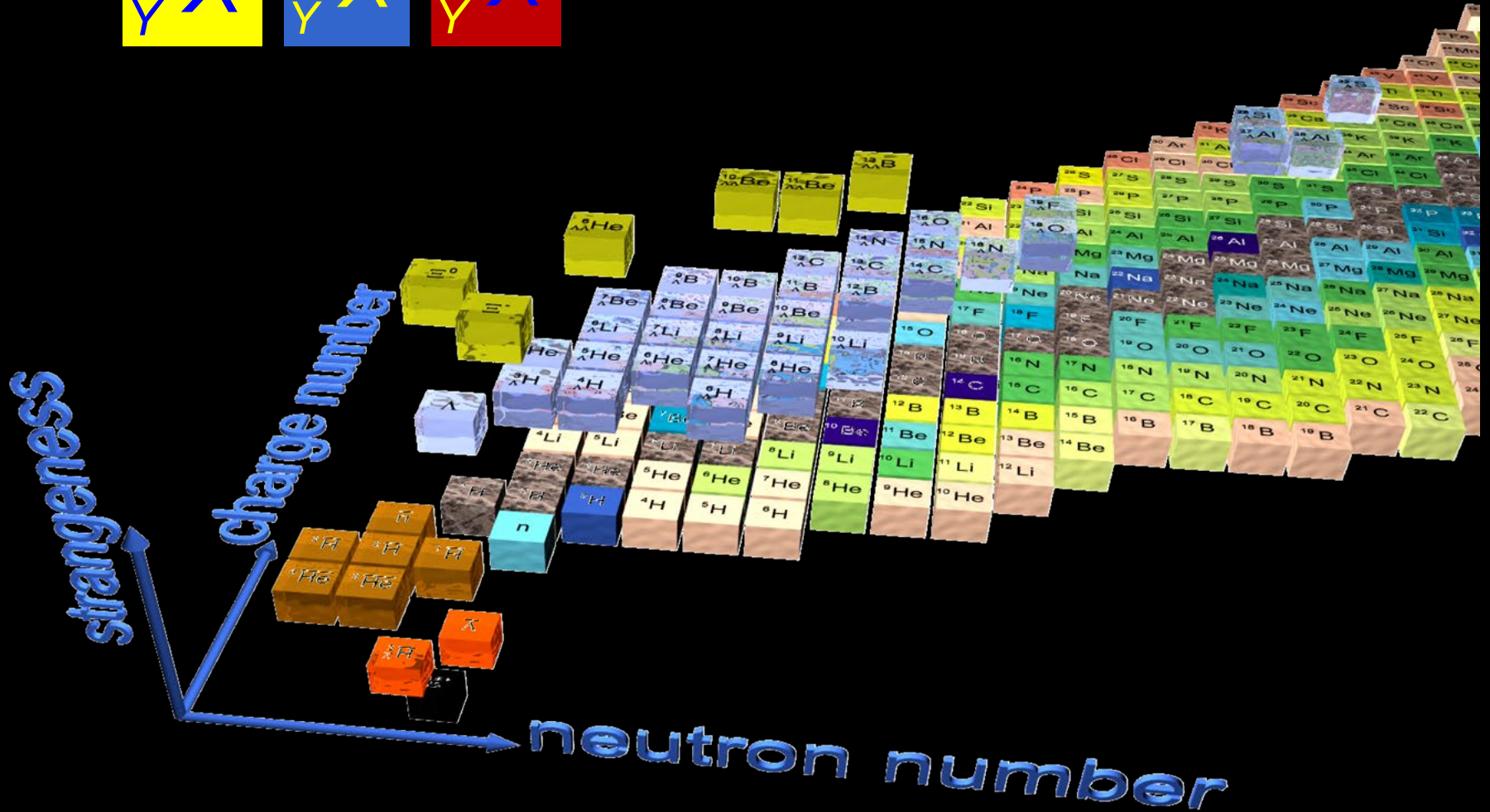
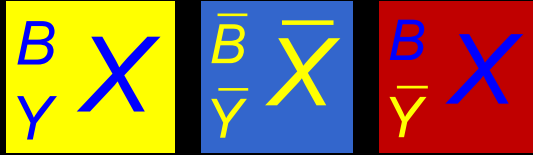
- **Examples**
- strangeness production $(\pi^+, K^+), (\pi^-, K^0)$
- strangeness exchange $(K^-, \pi^-), (K^-, \pi^0), (K^-, K^+)$
- electroproduction $(e, e^- K^+), (\gamma, K^+)$

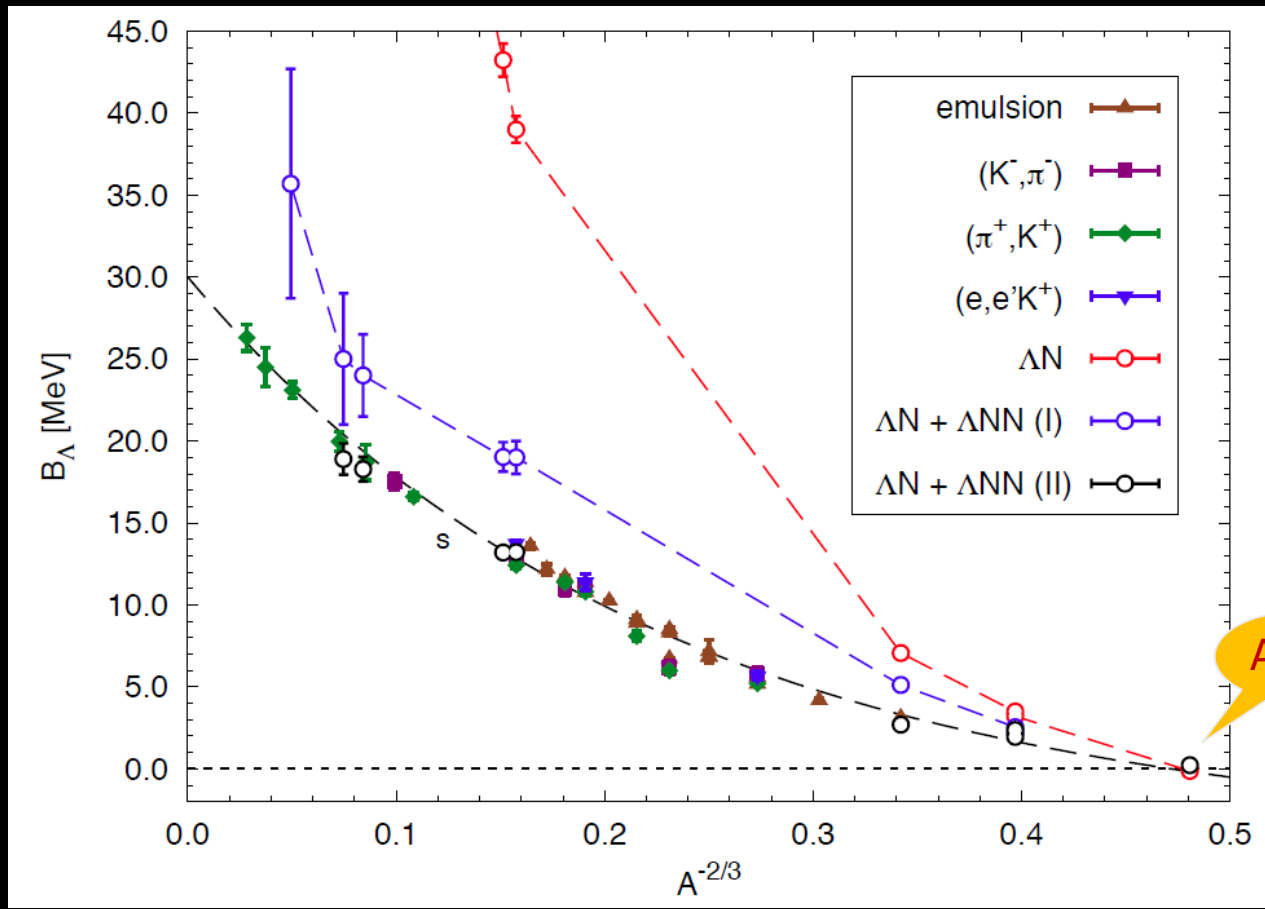
• **DECAY**

• **Examples**



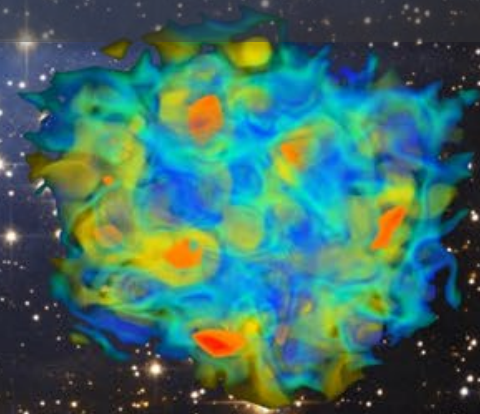
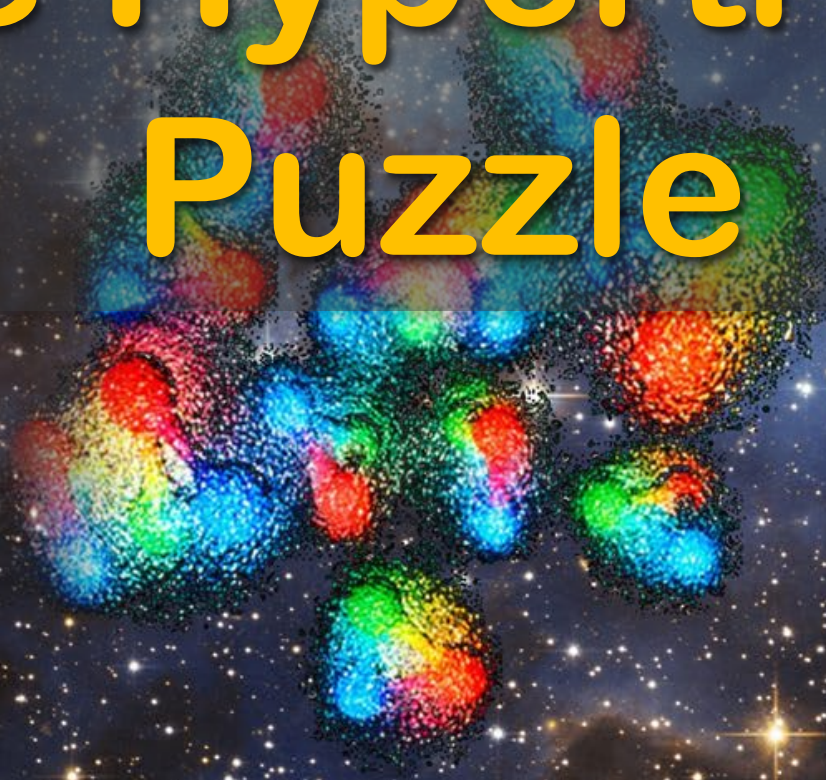
Hypernuclei are very exotic nuclei



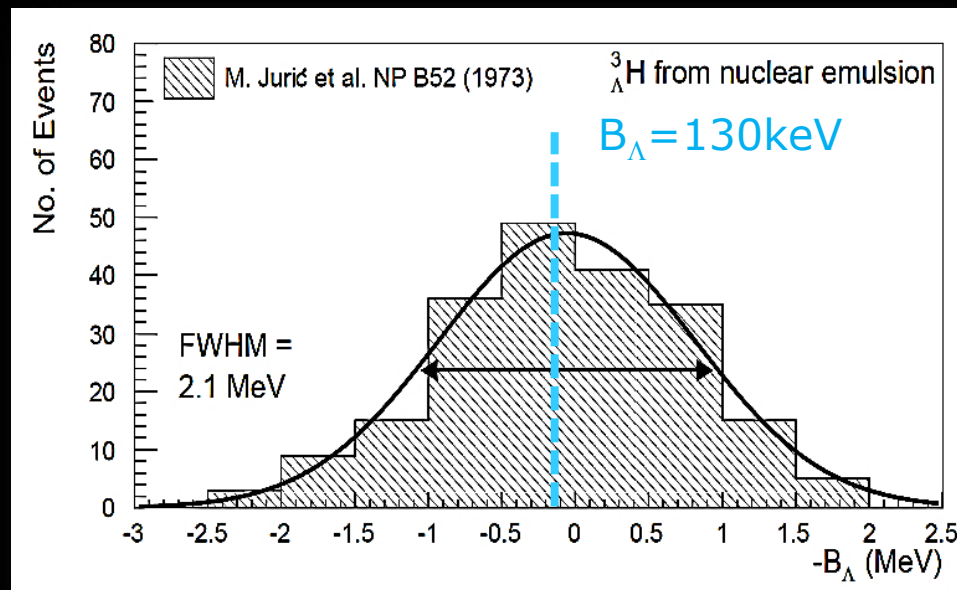
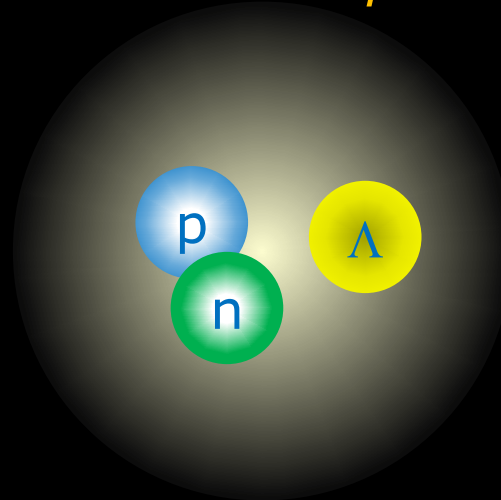


Three baryon interactions involving hyperons are essential

The Hypertriton Puzzle



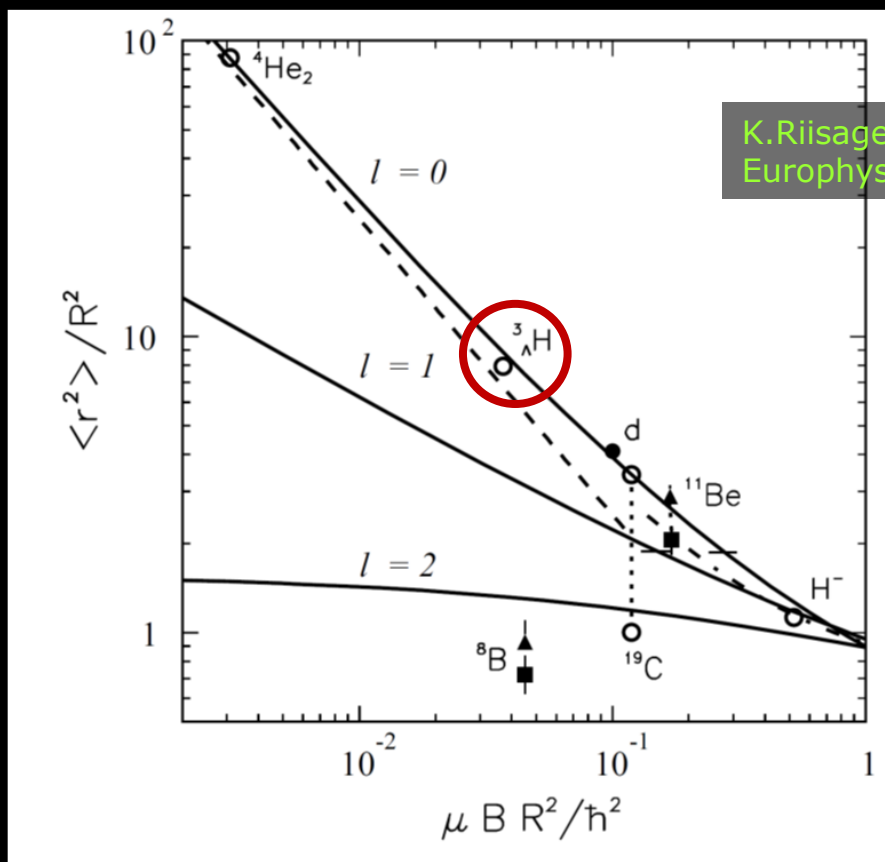
Do we understand the simplest Hypernucleus?



- $^3_{\Lambda}H$ is most fascinating halo nucleus
 - Binding energy $\approx 130\text{keV}$ \Rightarrow Characteristic length of two-body s-wave halo system small

$$\langle \Delta r^2 \rangle = \hbar^2 / (4\mu B) \xrightarrow{^3_{\Lambda}H} 10\text{ fm}$$

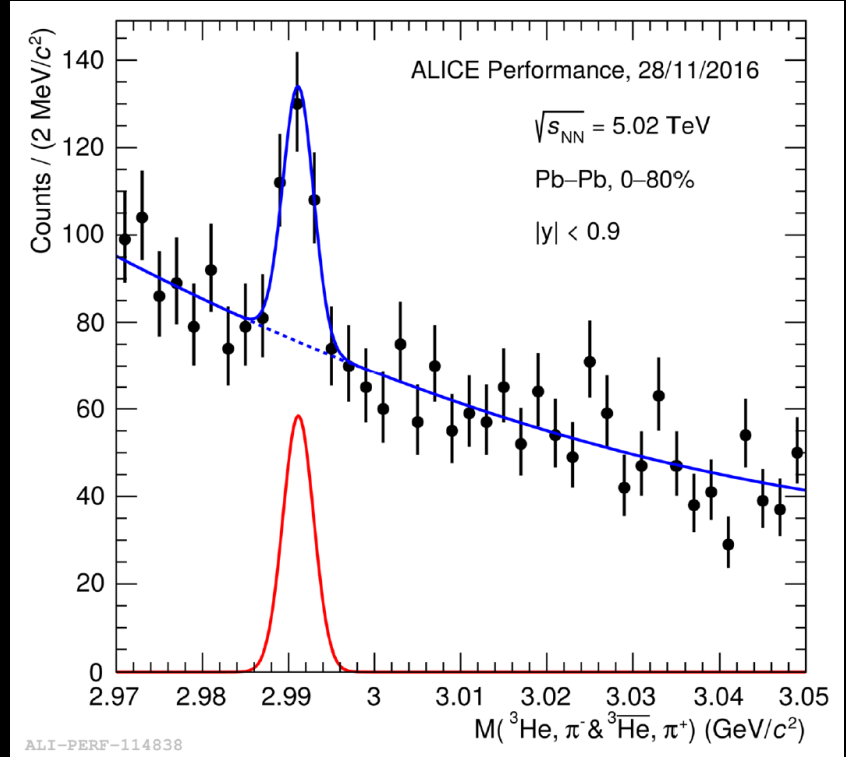
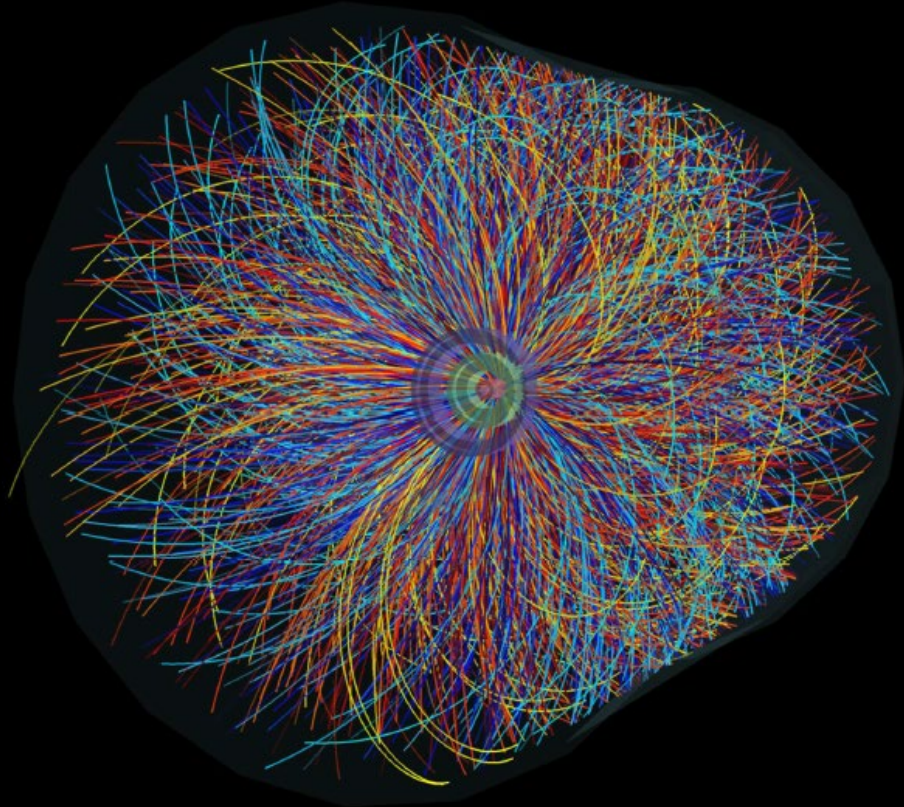
ratio of halo and core-potential square radii



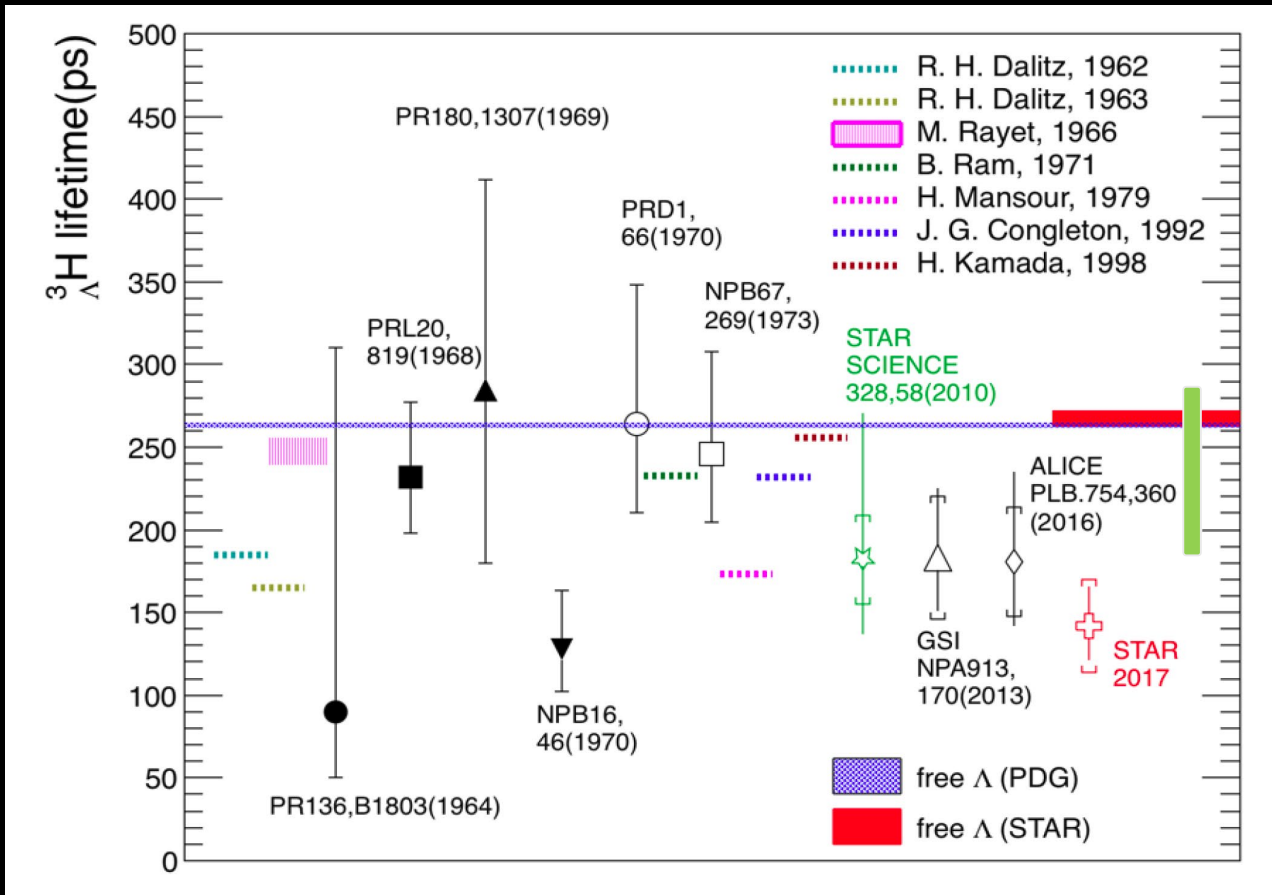
K. Riisager, D.V. Fedorov and A.S. Jensen, Europhys. Lett 49, 547 (2000)

scaled separation energy

➤ Searching the needle in the haystack



Experiment	Reaction	$\langle y/y_{cm} \rangle$	$\sqrt{s_{NN}}$ [GeV]	${}^3_{\Lambda}H$	$\overline{{}^3_{\Lambda}H}$	${}^4_{\Lambda}H$
E864	Au+Pt	0.3	5.0	1220 ± 854	-	-
HADES	Ar+KCl	-0.45	2.6	$\frac{{}^3H}{{N_{\Lambda}}} < 2.5 \cdot 10^{-2}$	-	-
STAR	Au+Au	0	7.7-200	≈ 400	≈ 200	-
ALICE	Pb+Pb	0	2760	≈ 124	≈ 90	-



ALICE, preliminary
 $237^{+33}_{-36}(\text{stat.}) \pm 17(\text{syst.}) \text{ps}$

STAR arXiv:1710.00436v1 [nucl-ex] 1st Oct 2017

small binding energy ? small lifetime

small binding energy

?

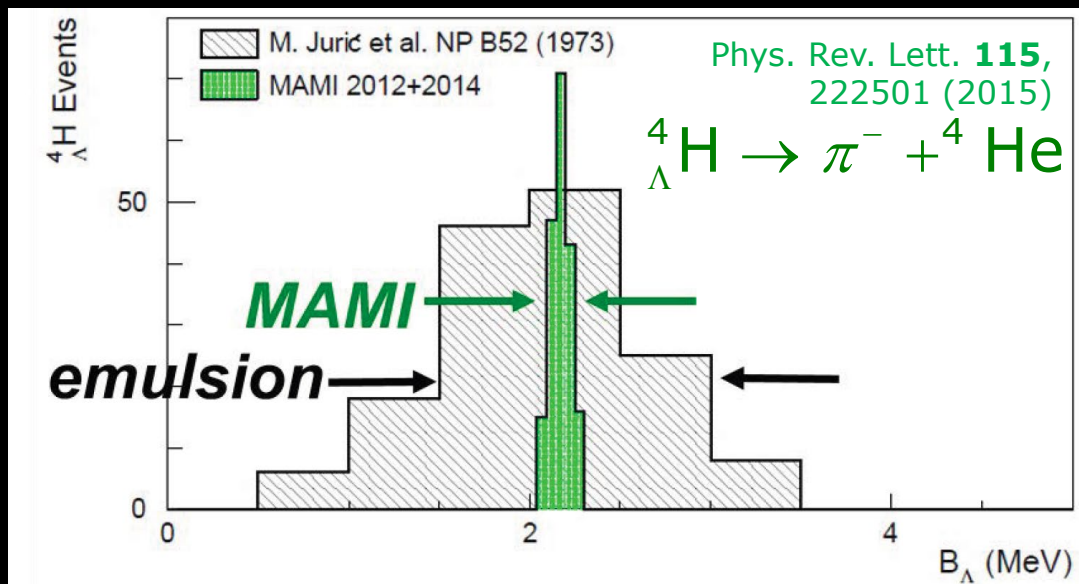
small lifetime

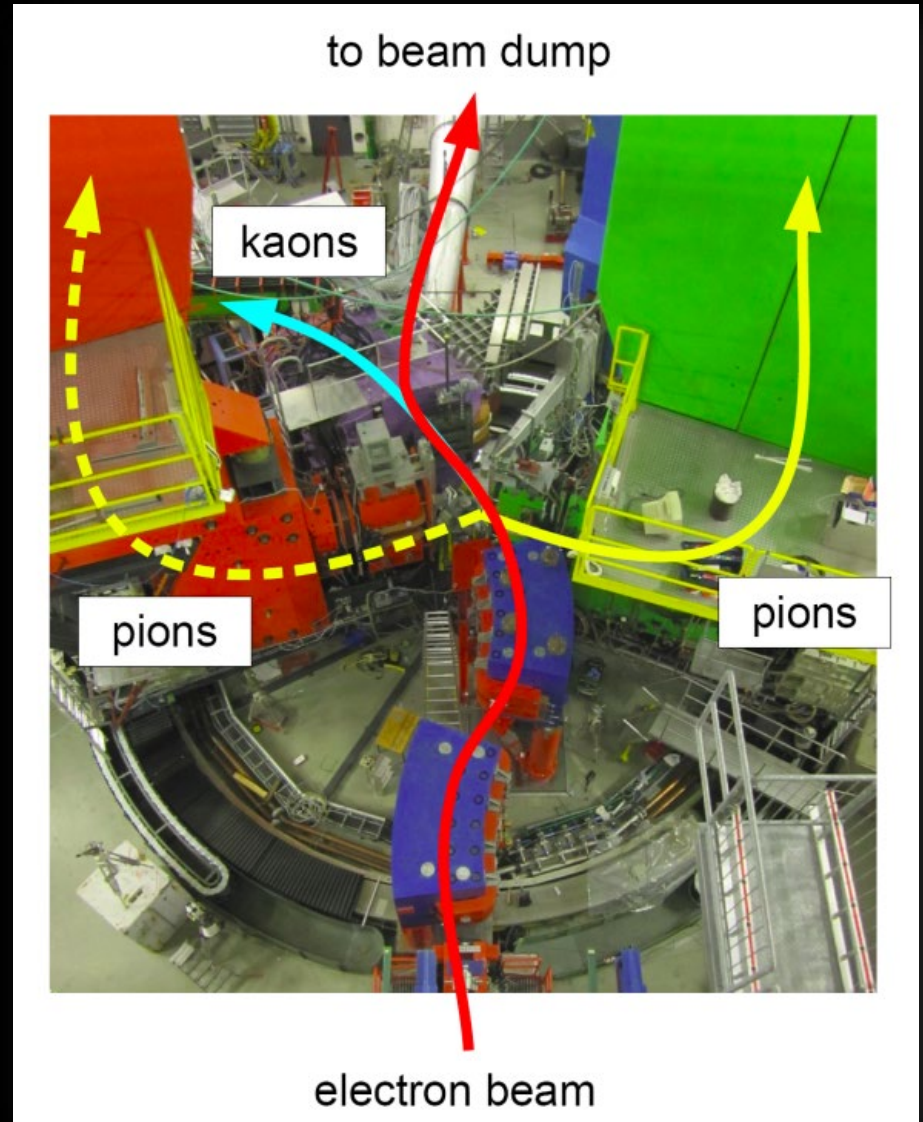
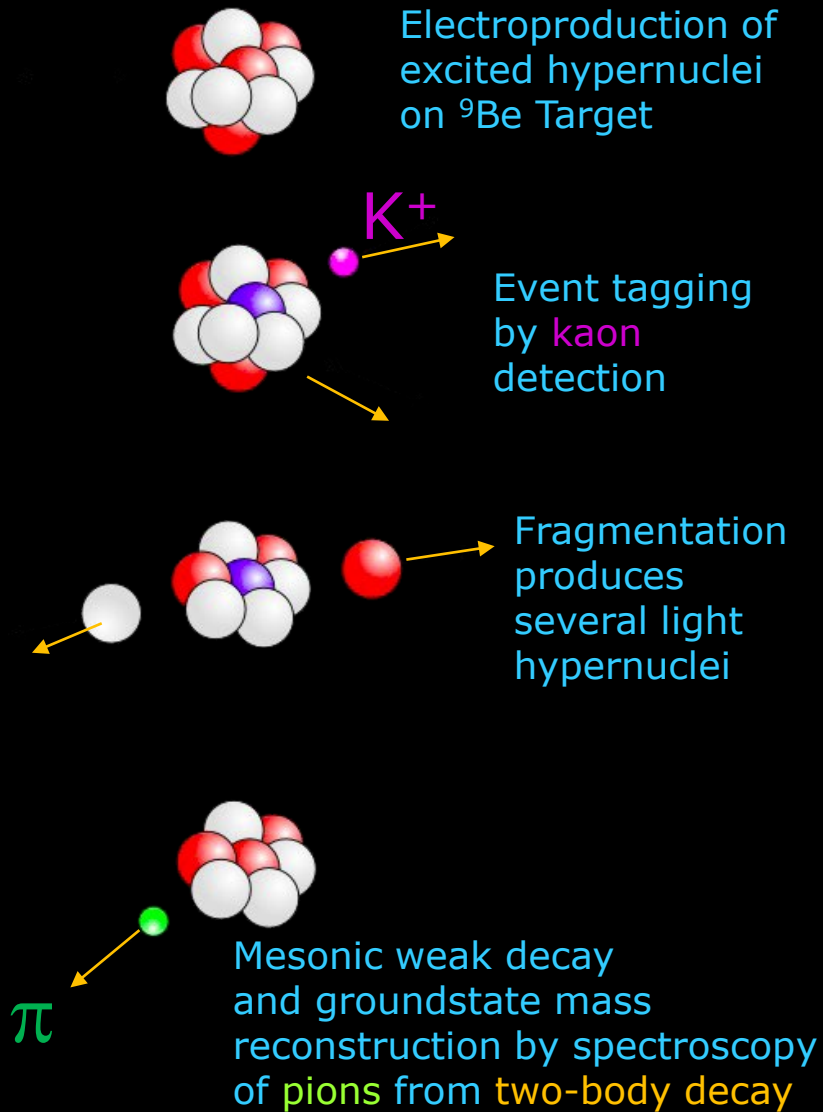
➤ New precision mass measurement at MAMI in 2020

- Make use of excellent beam quality at MAMI
- Precision *absolute* energy calibration interference of undulator radiation

➤ new lifetime measurements

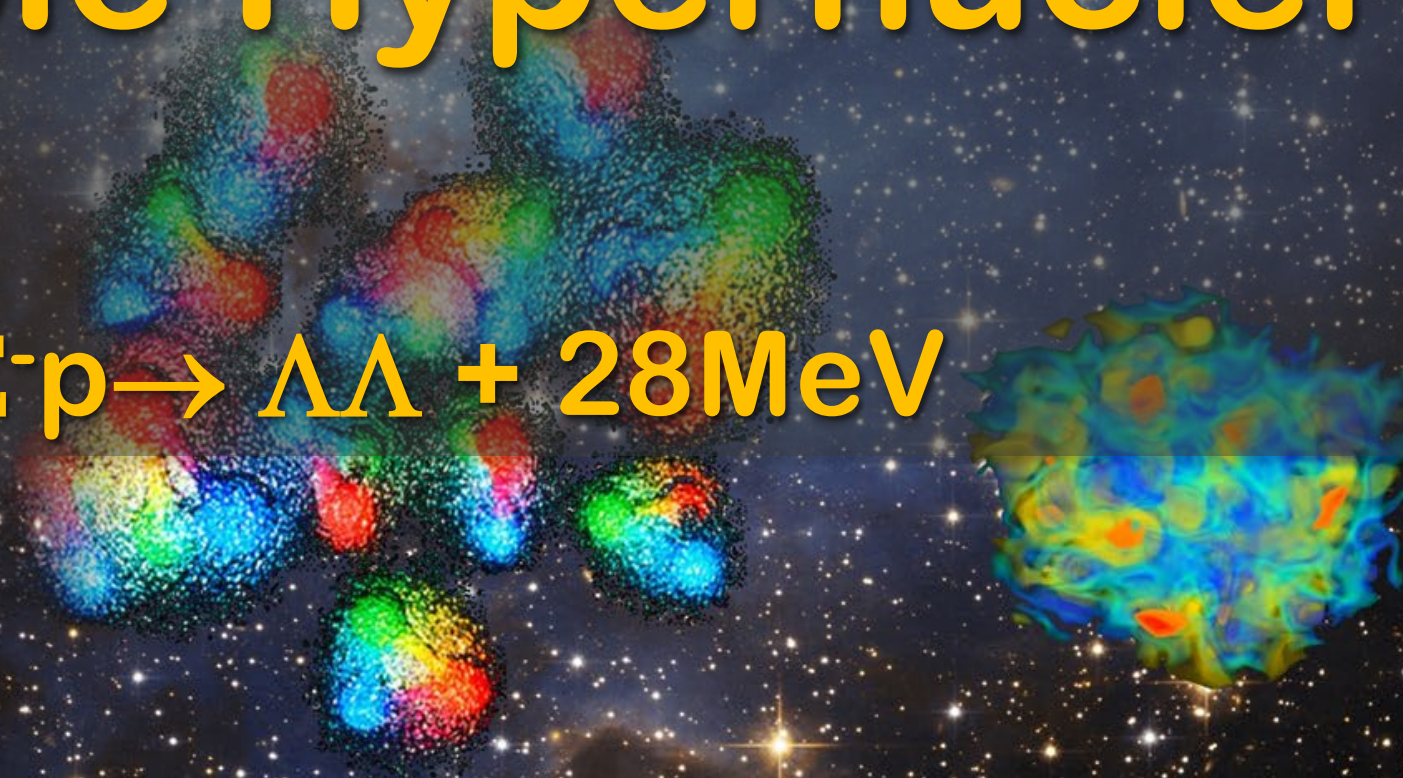
- 2020: ELPH (γ, K^+)
- 2020: WASA @ GSI/FAIR
- 2018: ALICE - end Run2: 2x statistics
- 2023: ALICE - end run 3: 200x stat.
- 202x: J-PARC (π^-, K^0)

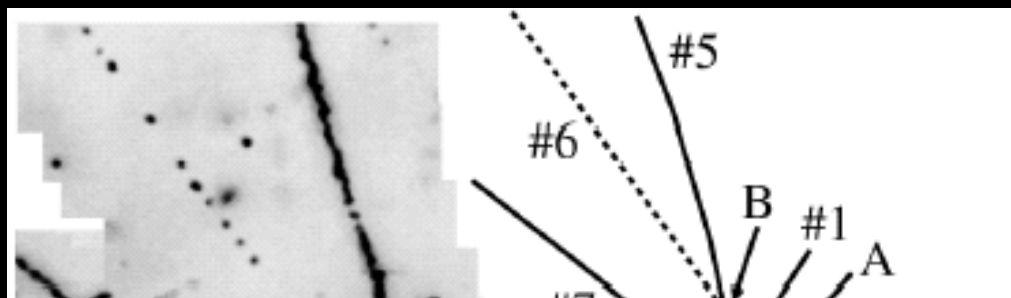




Double Hypernuclei

$$\Xi^- p \rightarrow \Lambda\Lambda + 28\text{MeV}$$





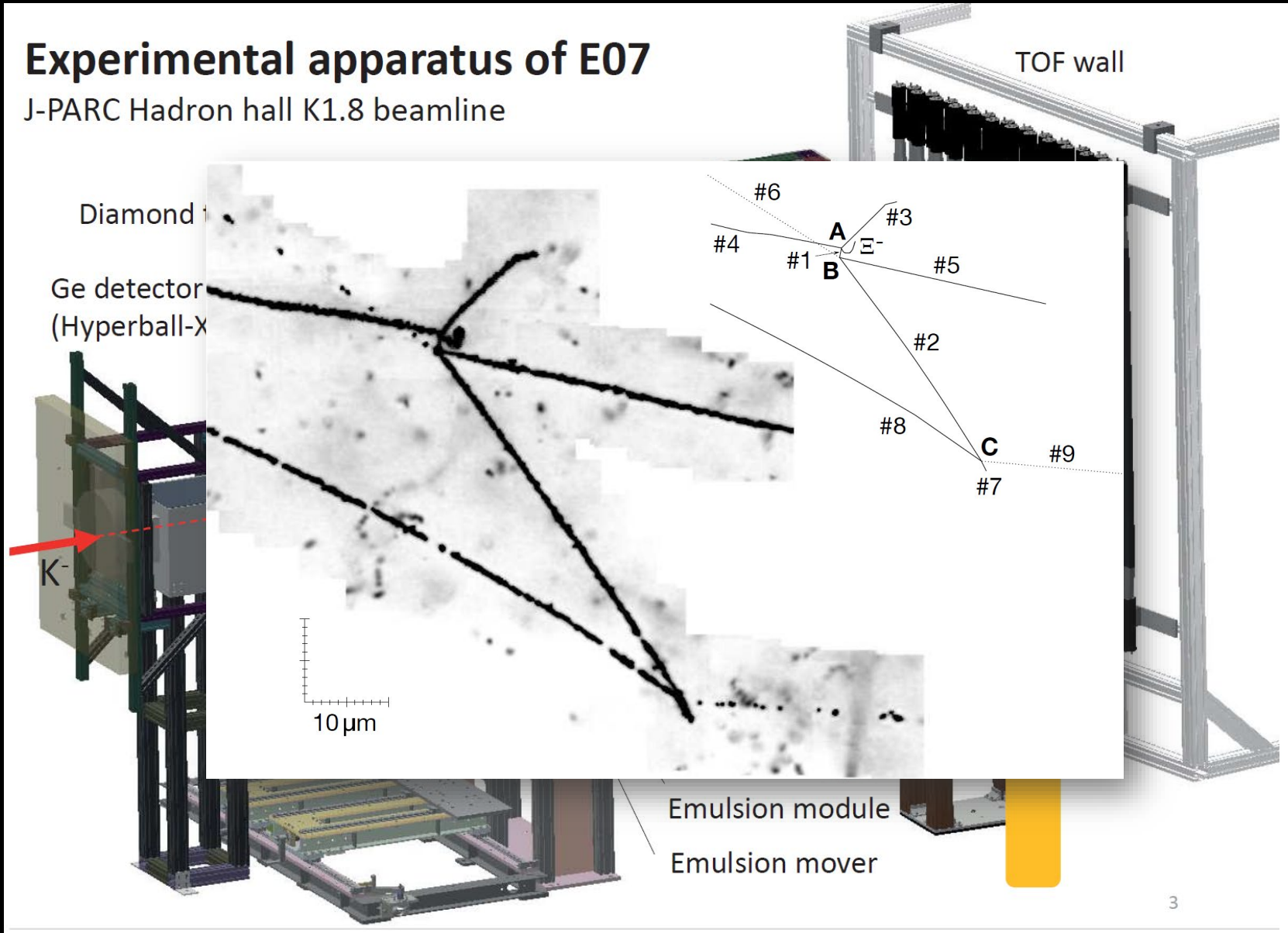
Nucleus	$\Delta B_{\Lambda\Lambda}(\Lambda\Lambda Z)$ [MeV]	Experiment	Reference	Remark
${}_{\Lambda\Lambda}^{10}\text{Be}$	4.3 ± 0.4	Danzysz (1963)	[179, 180] [174]	K^- + nuclear emulsion; $\Delta B_{\Lambda\Lambda}$ consistent with NAGARA if decay to ${}_{\Lambda}^9\text{Be}^*$ at $E_x \approx 3$ MeV [20, 181]
${}_{\Lambda\Lambda}^6\text{He}$	4.7 ± 0.6	Prowse (1966)	[475]	K^- + nuclear emulsion only schematic drawing
${}_{\Lambda\Lambda}^{10}\text{Be}$ or ${}_{\Lambda\Lambda}^{13}\text{B}$	-4.9 ± 0.7 0.6 ± 0.8	KEK-E176 (1991) Aoki event	[47, 618] [49, 195, 424]	hybrid-emulsion $(K^-, K^+)\Xi_{\text{stopped}}^-$
${}_{\Lambda\Lambda}^6\text{He}$	0.67 ± 0.17	KEK-E373 (2001) NAGARA event	[424, 557] [20]	hybrid emulsion
${}_{\Lambda\Lambda}^{10}\text{Be}$ or ${}_{\Lambda\Lambda}^{10}\text{Be}^*$	-1.65 ± 0.15	KEK-E373 (2001) DEMACHIYANAGI event	[21, 424] [20]	$B_{\Lambda\Lambda}$ consistent with Danzysz if $E_x \approx 2.8$ MeV
${}_{\Lambda\Lambda}^6\text{He}$ or ${}_{\Lambda\Lambda}^{11}\text{Be}^*$	3.77 ± 1.71 3.95 ± 3.00 or 4.85 ± 2.63	KEK-E373 (2003) MIKAGE event	[20, 558]	
${}_{\Lambda\Lambda}^{12}\text{Be}$ or ${}_{\Lambda\Lambda}^{11}\text{Be}^*$	2.00 ± 1.21 2.61 ± 1.34	KEK-E373 (2010) HIDA event	[20, 424]	
${}_{\Lambda\Lambda}^{10}\text{Be}$ or ${}_{\Lambda\Lambda}^{11}\text{Be}$ or ${}_{\Lambda\Lambda}^{12}\text{Be}$	1.63 ± 0.14 1.87 ± 0.37 -2.7 ± 1.0	J-PARC E07 MINO event	[349]	most probable ${}_{\Lambda\Lambda}^{11}\text{Be}$

randomness

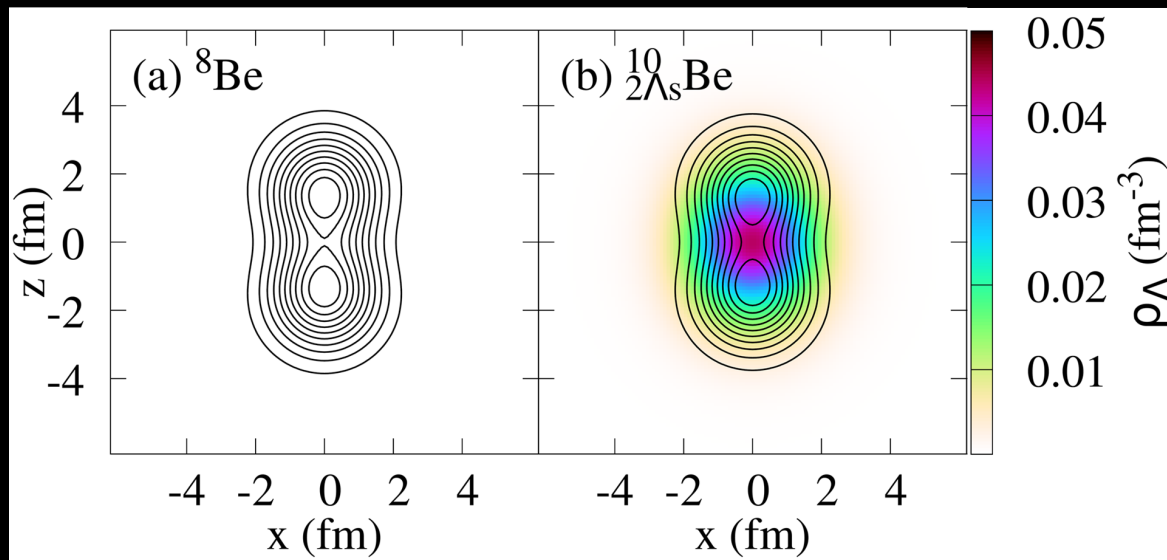
number

Experimental apparatus of E07

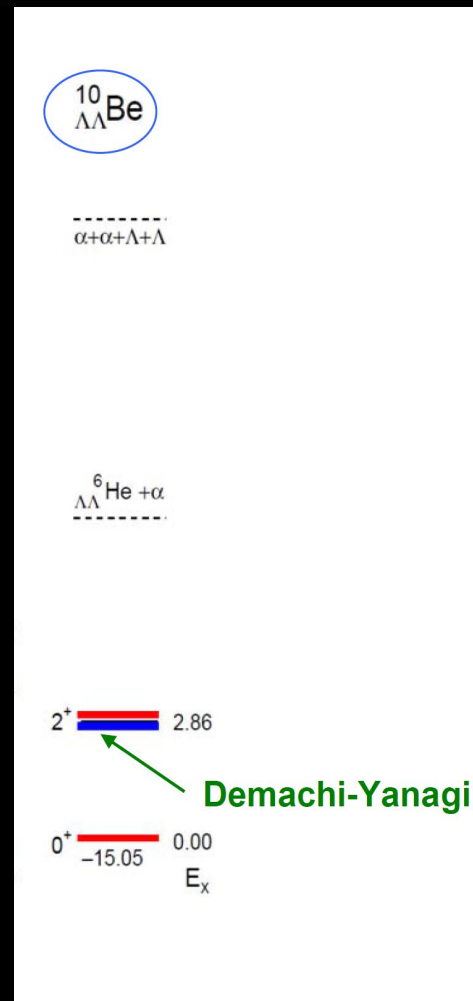
J-PARC Hadron hall K1.8 beamline



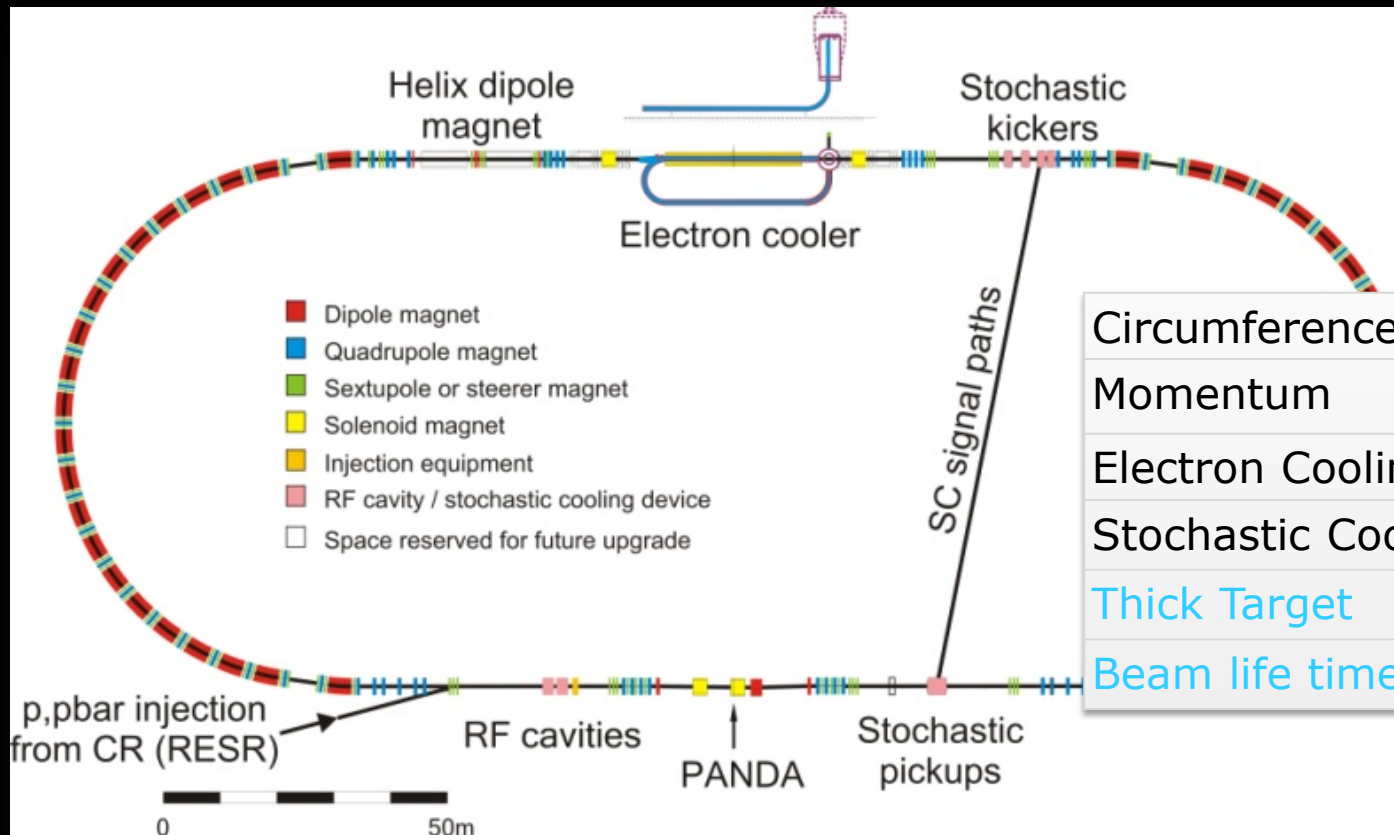
E. Hiyama, M. Kamimura, T. Motoba, T. Yamada and Y. Yamamoto
Phys. Rev. 66 (2002), 024007



Yusuke Tanimura
Phys. Rev. C 99, 034324 (2019)



- many excited, particle stable states in double hypernuclei predicted
- level structure reflects in 0th order levels of core nucleus



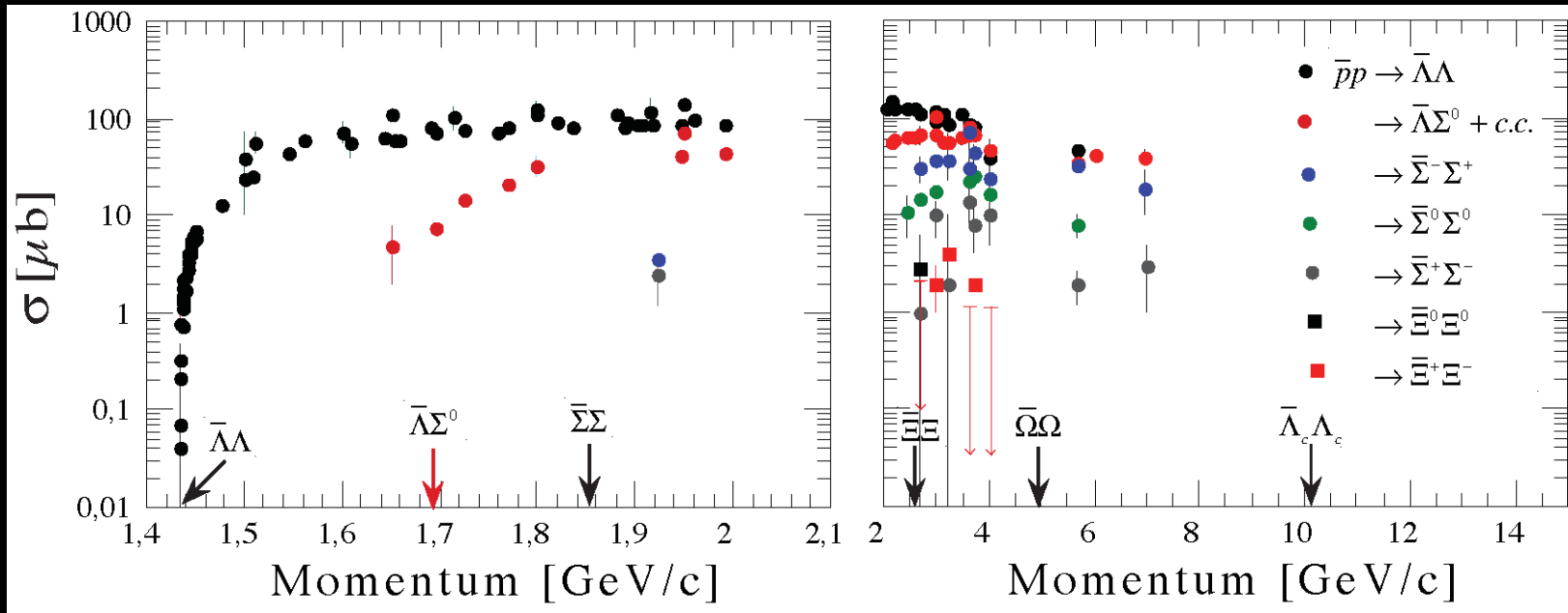
Circumference	575 m
Momentum	1.5 – 15 GeV/c
Electron Cooling	up to 9 GeV/c
Stochastic Cool.	Full range
Thick Target	$4 \cdot 10^{15} \text{ cm}^{-2}$
Beam life time	>30 min

➤ High resolution mode

- e^- cooling $1.5 \leq p \leq 8.9 \text{ GeV/c}$
- 10^{10} antiprotons stored
- Luminosity up to $2 \cdot 10^{31} \text{ cm}^{-2}\text{s}^{-1}$
- $\Delta p/p \leq 4 \cdot 10^{-5}$

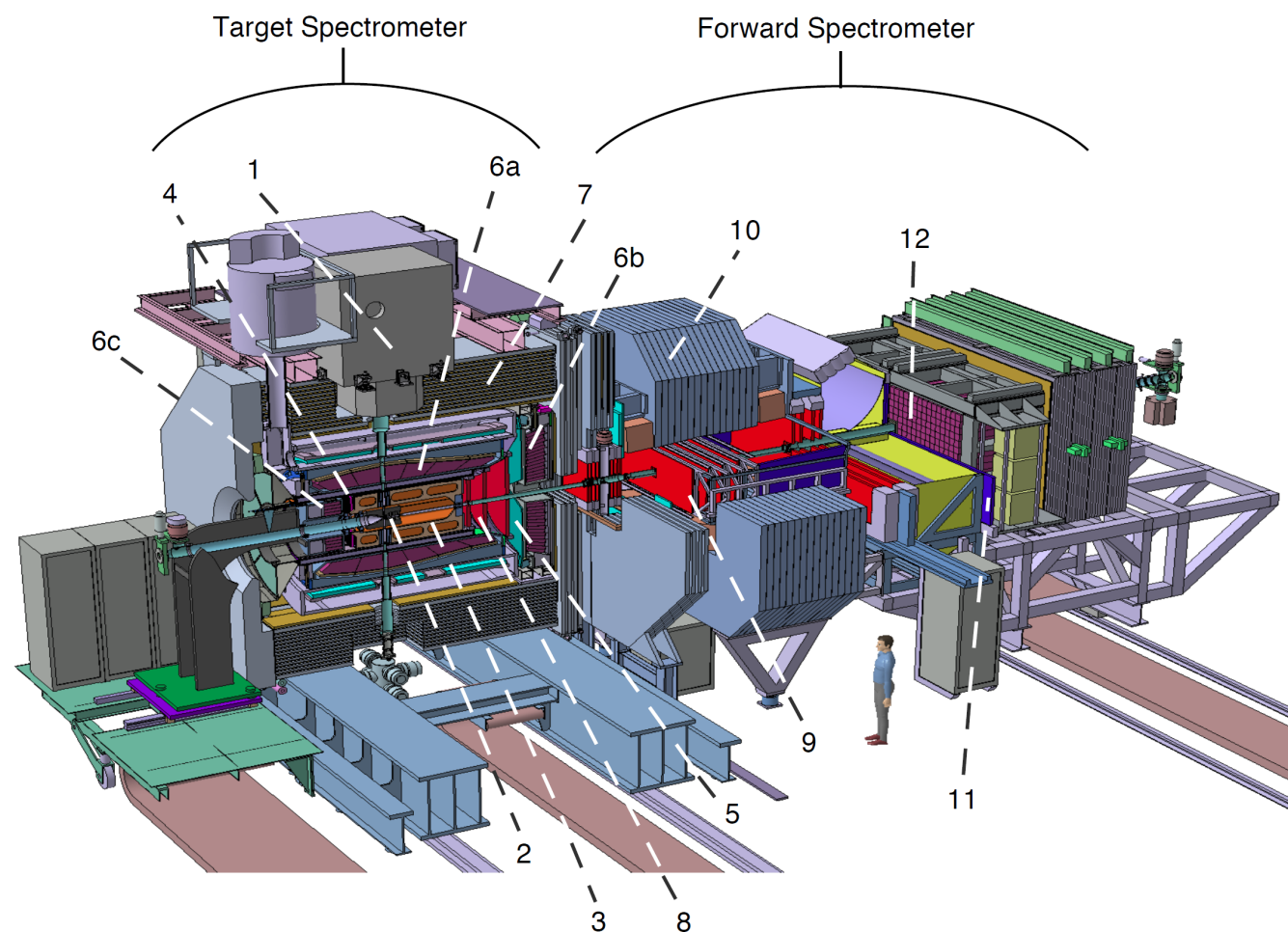
➤ High luminosity mode

- Stochastic cooling $p \geq 3.8 \text{ GeV/c}$
- 10^{11} antiprotons stored
- Luminosity up to $2 \cdot 10^{32} \text{ cm}^{-2}\text{s}^{-1}$
- $\Delta p/p \leq 2 \cdot 10^{-4}$

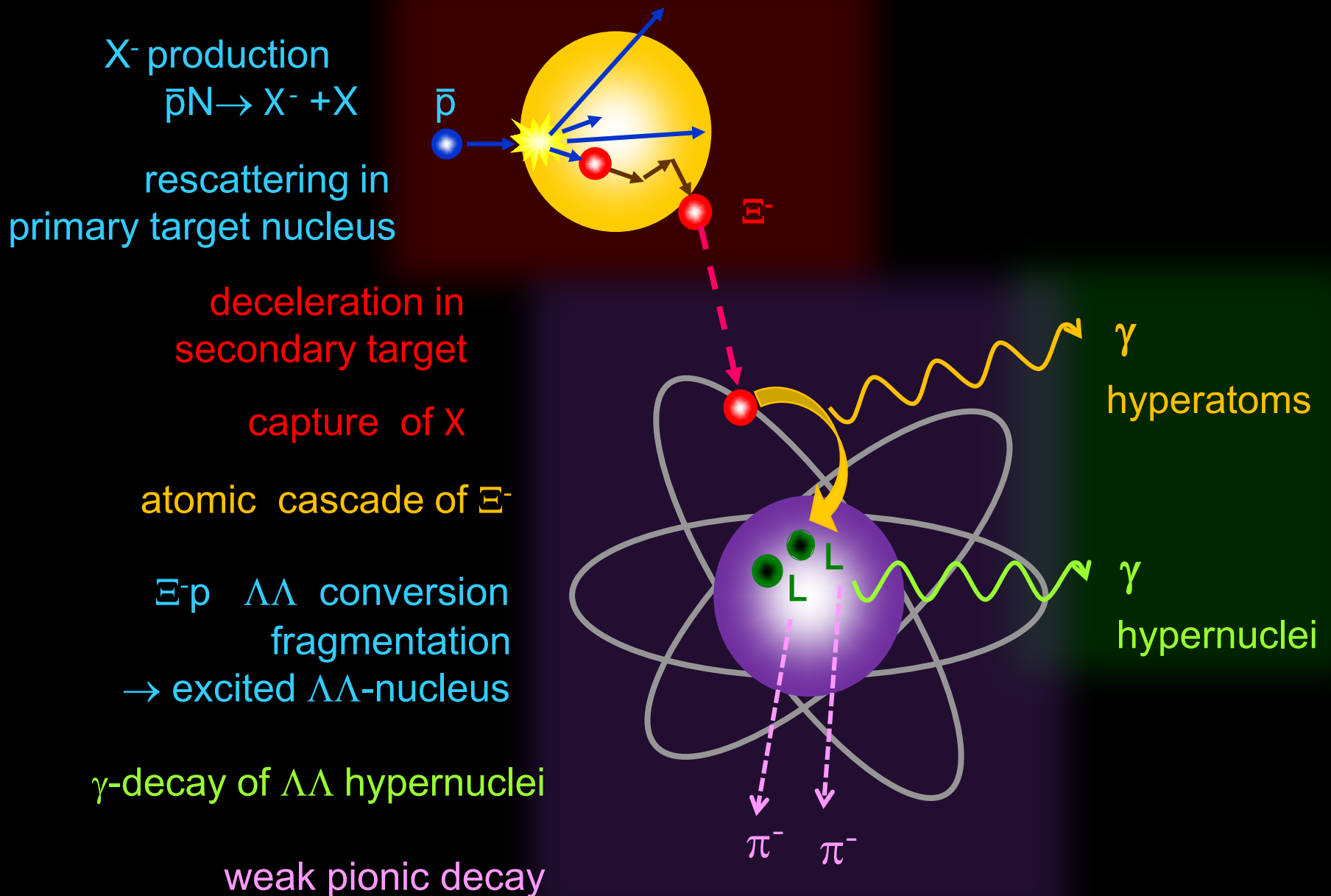


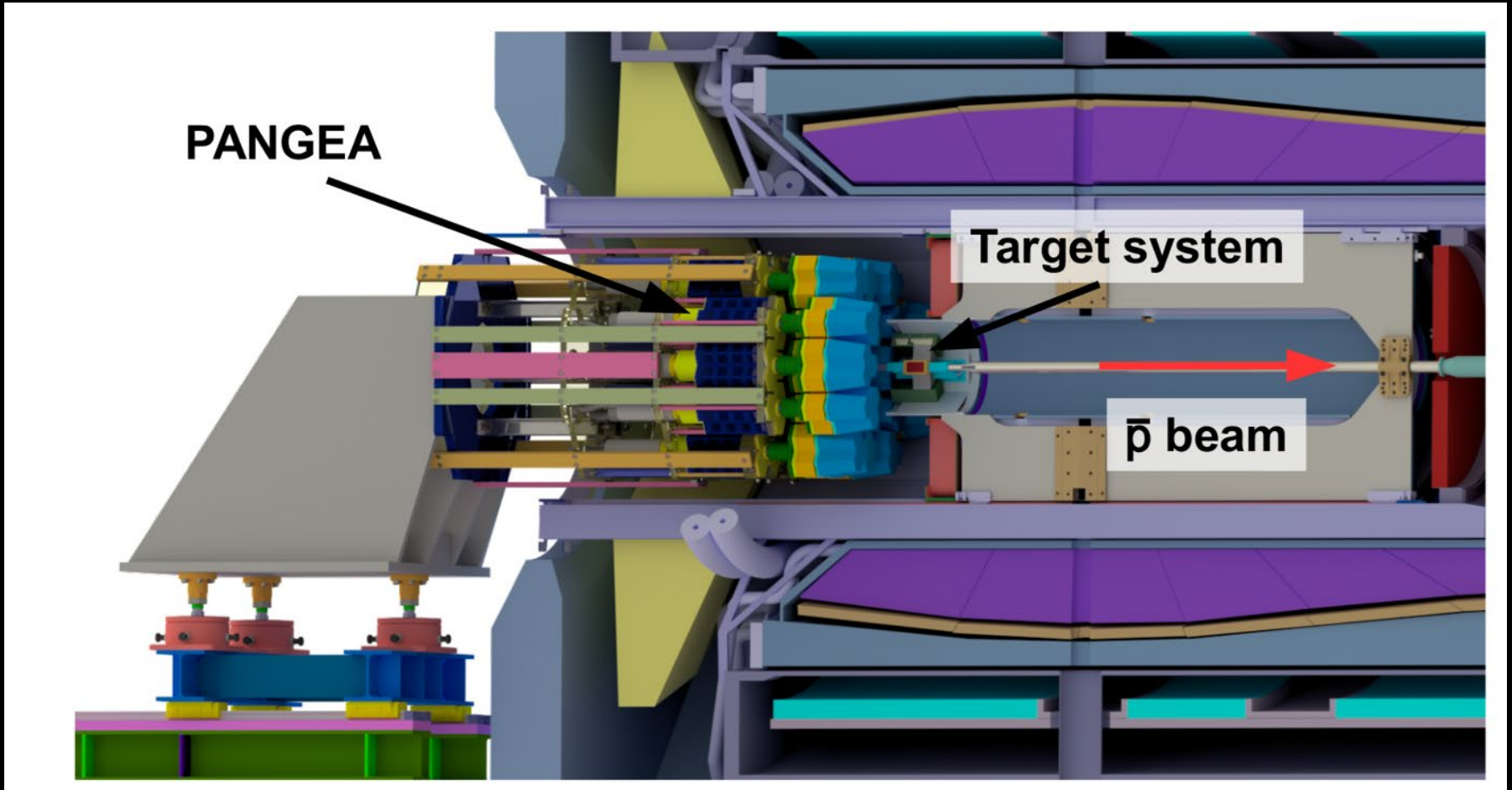
Production Rates (1-2 (fb)⁻¹/y)

<u>Final State</u>	<u>cross section</u>	<u># reconstr. events/y</u>
Meson resonance + anything	100μb	10 ¹⁰
$\Lambda\bar{\Lambda}$	50μb	10 ¹⁰
$\Xi\bar{\Xi} (\rightarrow \Lambda\Lambda A)$	2μb	10 ⁸ (10 ⁵)
$D\bar{D}$	250nb	10 ⁷
$J/\psi (\rightarrow e^+e^-, \mu^+\mu^-)$	630nb	10 ⁹
$\chi_2 (\rightarrow J/\psi + \gamma)$	3.7nb	10 ⁷
$\Lambda_c\bar{\Lambda}_c$	20nb	10 ⁷
$\Omega_c\bar{\Omega}_c$	0.1nb	10 ⁵



- | | | |
|---|-------------------------------------|----------------------|
| 1: Cluster-Target | 6a: Barrel-EMC | 9: Forward Tracker |
| 2: Mikrovertex-Detektor | 6b: Forward-EMC | 10: Dipole |
| 3: STT-Tracker | 6c: Backward-EMC | 11: Forward TOF wall |
| 4: DIRC | 7: Solenoid Yoke with Muon Chambers | 12: Shashlyk-EMC |
| 5: Disc-DIRC | 8: GEM-Tracker | |
| not visible (downstream): Luminosity Detector | | |
| not visible: Hypernuclear Setup | | |







- Tools**
- heavy ion beams
 - electron beams
 - photon beams
 - meson beams
 - antiproton beams

- Methods**
- missing mass studies
 - invariant mass studies
 - γ -spectroscopy
 - π -spectroscopy
 - FSI

- Observables**
- masses
 - excitation spectrum
 - lifetimes
 - branching ratios
 - cross section

Take-home message

- Strangeness nuclear physics is embedded in the quest to determine the EOS of dense stellar systems
- Hyperon puzzle of neutron stars is still not solved
- Hypernuclei and hyperatoms are femto-laboratories for $Y^n N^m$ interaction
- After 60 years still many puzzles: hypertriton, existence of neutral hypernuclei $nn\Lambda$, $nn\Lambda\Lambda$, ...
- Coming generation of experiments focus on precision studies



**Thank you
for your attention**